

# MiniBooNE $\nu_\mu$ and $\bar{\nu}_\mu$ disappearance results

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# Overview

- 1) Neutrino oscillation
- 2) MiniBooNE experiment
- 3) MiniBooNE-only neutrino disappearance analysis
- 4) Antineutrino disappearance analysis
- 5) Improvements to disappearance analysis
- 6) Conclusion

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# Neutrino oscillation

Neutrinos “oscillate” because the flavor state of the neutrino,  $\nu_\alpha$ , is related to the mass states,  $\nu_i$ , by a mixing matrix,  $U_{\alpha i}$

$$|\nu_i\rangle = \sum U_{\alpha i} |\nu_\alpha\rangle$$

Since there are three observed flavors of neutrinos ( $\nu_e, \nu_\mu, \nu_\tau$ ), the mixing matrix,  $U_{\alpha i}$ , contains three mixing angles ( $\theta_{12}, \theta_{23}, \theta_{13}$ ) and a CP violating phase  $\delta$ . It can be factorized into three blocks, each corresponding to two neutrino mixing.

$$U_{\alpha i} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$c_{ij} = \cos\theta_{ij}, \quad s_{ij} = \sin\theta_{ij}$$

# Neutrino oscillation

As the states propagate in time, the neutrino mass states interfere:

$$| \underline{v}_\alpha(t) \rangle = \sum -\sin\theta_{ij} | \underline{v}_i \rangle + \cos\theta_{ij} | \underline{v}_j \rangle$$

The probability to observe  $v_\beta$  with a pure  $v_\alpha$  sample is:

$$P_{\alpha \rightarrow \beta} = |\langle v_\beta | v_\alpha(t) \rangle|^2 = \sin^2 2\theta_{ij} \sin^2 \left( 1.27 \frac{\Delta m_{ij}^2 L}{E} \right)$$

where  $L$  (km) is the distance traveled,  $E$  (GeV) is the energy of the neutrino and  $\Delta m^2$  (eV<sup>2</sup>) is the difference of the masses squared:

$$\Delta m_{ij}^2 = m_i^2 - m_j^2$$

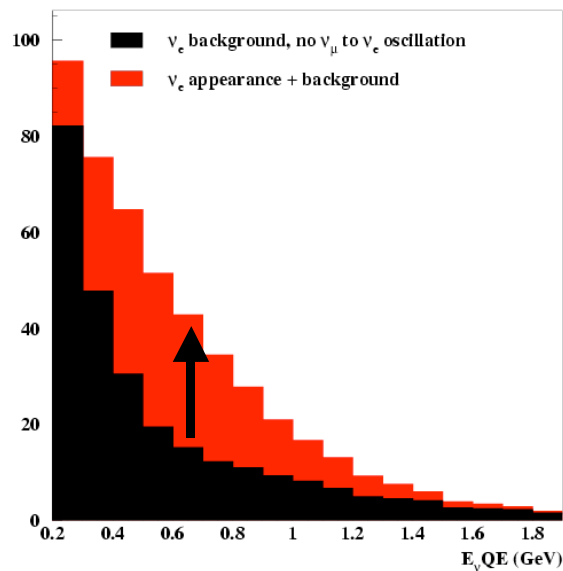
Choice of  $L$  and  $E$  chooses what range of  $\Delta m^2$  the experiment is sensitive to, the size of the oscillations sets  $\sin^2 2\theta$   
3 neutrino masses mean 2 independant  $\Delta m^2$

# Disappearance and Appearance experiments

Starting with a  $\nu_\alpha$  beam, there are two ways to look for oscillation:

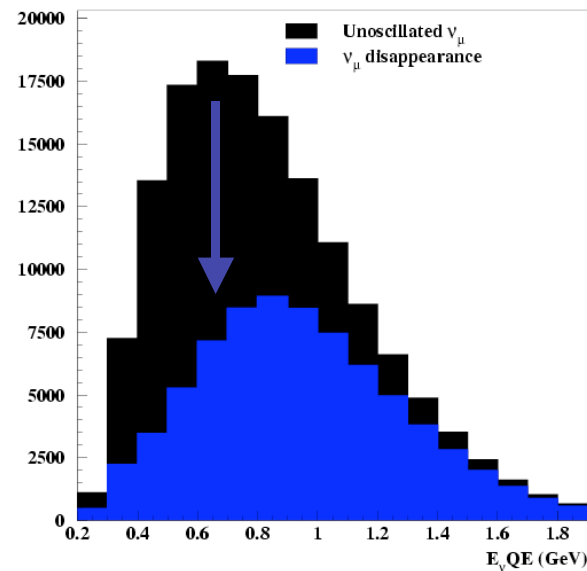
**Appearance** experiment

Detect more  $\nu_\beta$  than expected



**Disappearance** experiment

Detect less  $\nu_\alpha$  than expected



Neutrinos at energy  $E_1$  oscillate differently than at  $E_2$  for the same  $L$ , creating a **unique signature for oscillation vs energy**

## Reducing errors with a second detector

Source of error	Total fractional error (%)
pBe $\rightarrow$ $\pi^+$ production (flux)	4.0
beamline and horn model (flux)	4.3
cross sections	18.6
detector model	4.0
total	19.9

Adding a second detector measures the flux x cross section to the level of uncorrelated errors between the two detectors

Start with 20% error

Remove flux, cross section, and beam errors: 20%  $\rightarrow$  4%

Add 5% uncorrelated errors: 4% + 5% = 6%

# Normalization information

To search for disappearance, can use **normalization** or **shape** information

## 1) Normalization information:

Compare total number of events to expectation  
(aka “counting experiment”)

K2K **expected**:  $158 + 9.2 - 8.6$  events at the far detector  
but **observed**: 112 events

Normalization information provided by additional detectors  
Limited by statistics at far detector

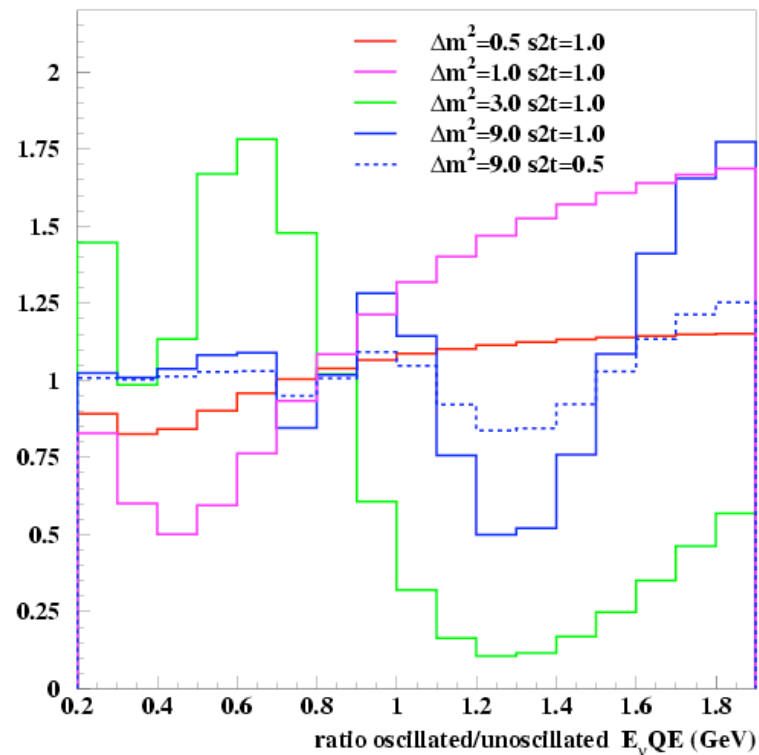


# Shape information

To search for disappearance, can use **normalization** or **shape** information

## 2) Shape information:

Compare the energy distribution of events to no oscillation hypothesis



Ratio of oscillated events/ unoscillated events vs energy

- $\Delta m^2$  changes the periodicity of the oscillation (see  $\Delta m^2=1\text{eV}^2$ ,  $\Delta m^2=3\text{eV}^2$ )
- $\sin^2 2\theta$  changes the depth of the oscillation (see  $\sin^2 2\theta=1.0$ ,  $\sin^2 2\theta=0.5$ )

MiniBooNE will make a one detector shape measurement

# Oscillation observations

Plot of all oscillation experiments:

“Solar”:  $\Delta m_{12}^2 \sim 10^{-5} \text{eV}^2$ ,  $\sin^2 2\theta_{12} \sim 32^\circ$

With solar  $\nu$ : SNO

With reactor  $\nu$ : KamLAND

“Atmospheric”:  $\Delta m_{23}^2 \sim 10^{-3} \text{eV}^2$ ,  
 $\sin^2 2\theta_{23} \sim 45^\circ$

With atmospheric  $\nu$ : SuperK

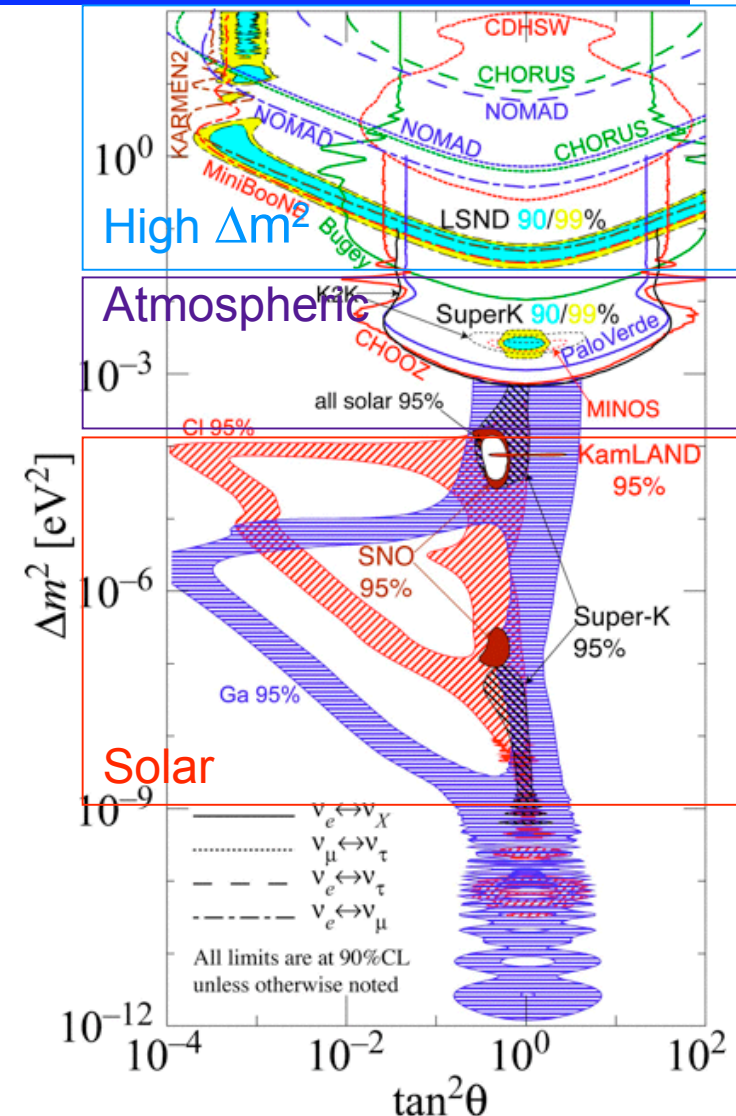
With Accelerator  $\nu$ : MINOS

“High  $\Delta m^2$ ”:  $\Delta m^2 \sim 1-10 \text{eV}^2$

CDHS (disappearance)

CCFR (disappearance)

LSND (appearance)

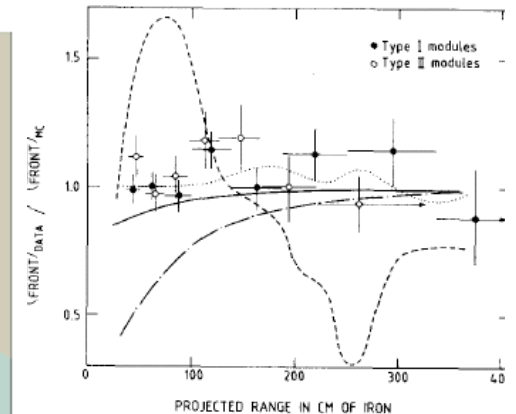
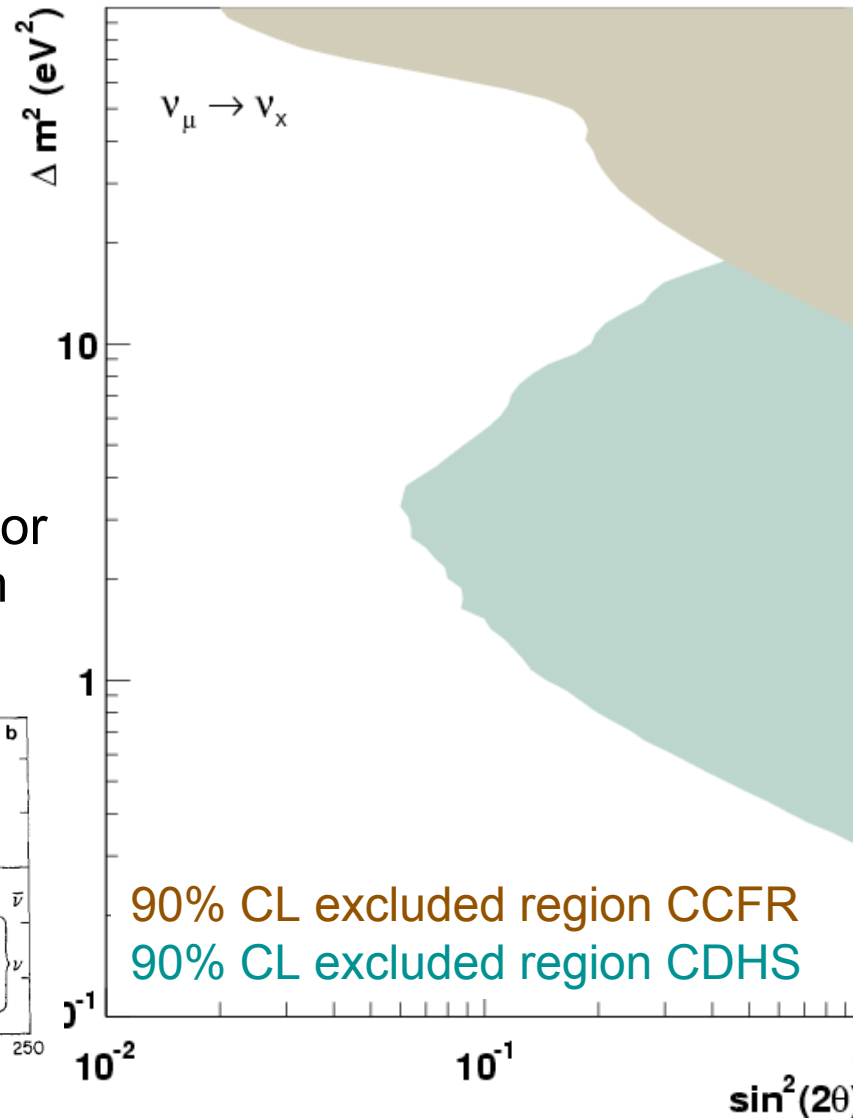
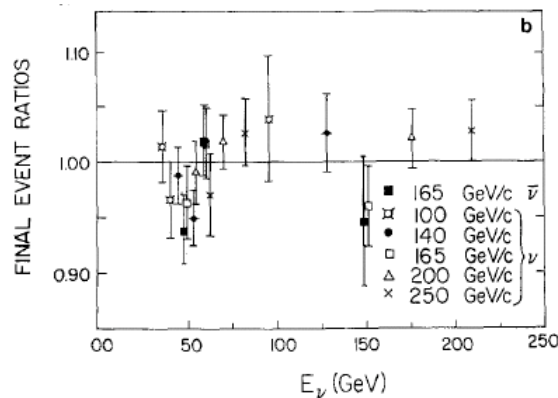


# High $\Delta m^2$ disappearance expts

## CCFR (FNAL E701)

I.E. Stockdale et al  
Z.Phys.C27:53,1985

- Mono energetic meson beam produces dichromatic ( $\sim 50, 160\text{GeV}$ ) neutrino beam
- Two steel/scintillator detectors at 715m and 1116m



## CDHS at CERN

F. Dydak et al.  
Phys.Lett.B134:281,1984.

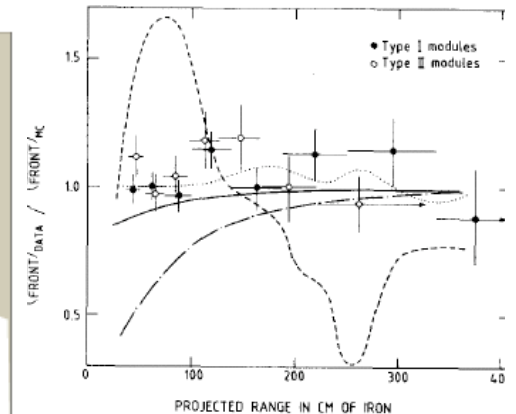
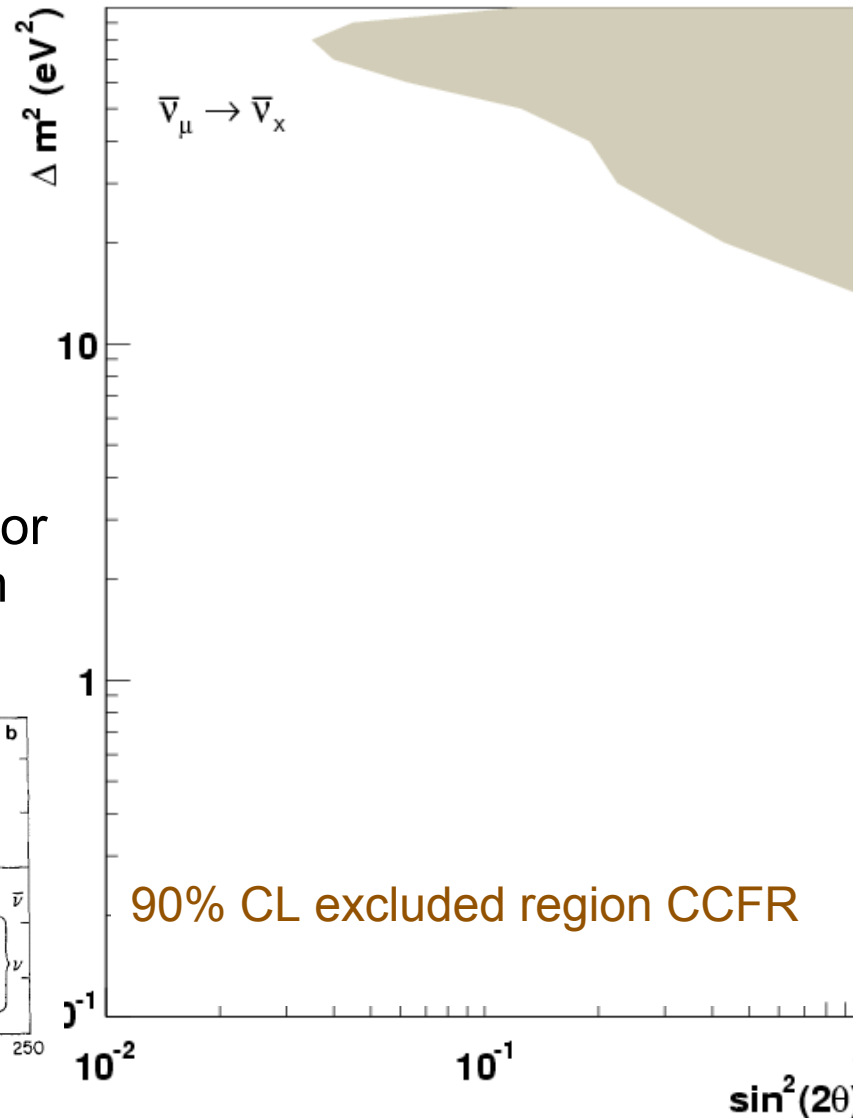
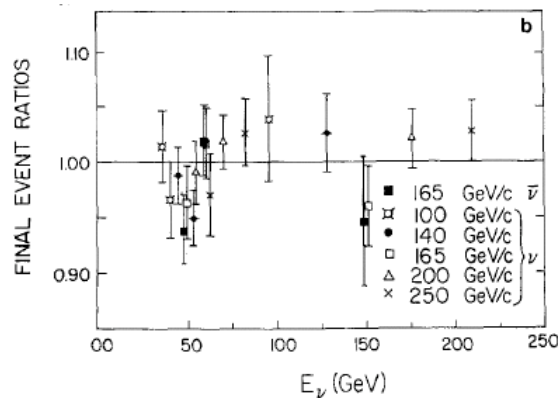
- 19.2 GeV protons on Be target produces  $\sim 3\text{GeV}$  neutrino beam
- Two iron/scintillator detectors at 130m and 885m

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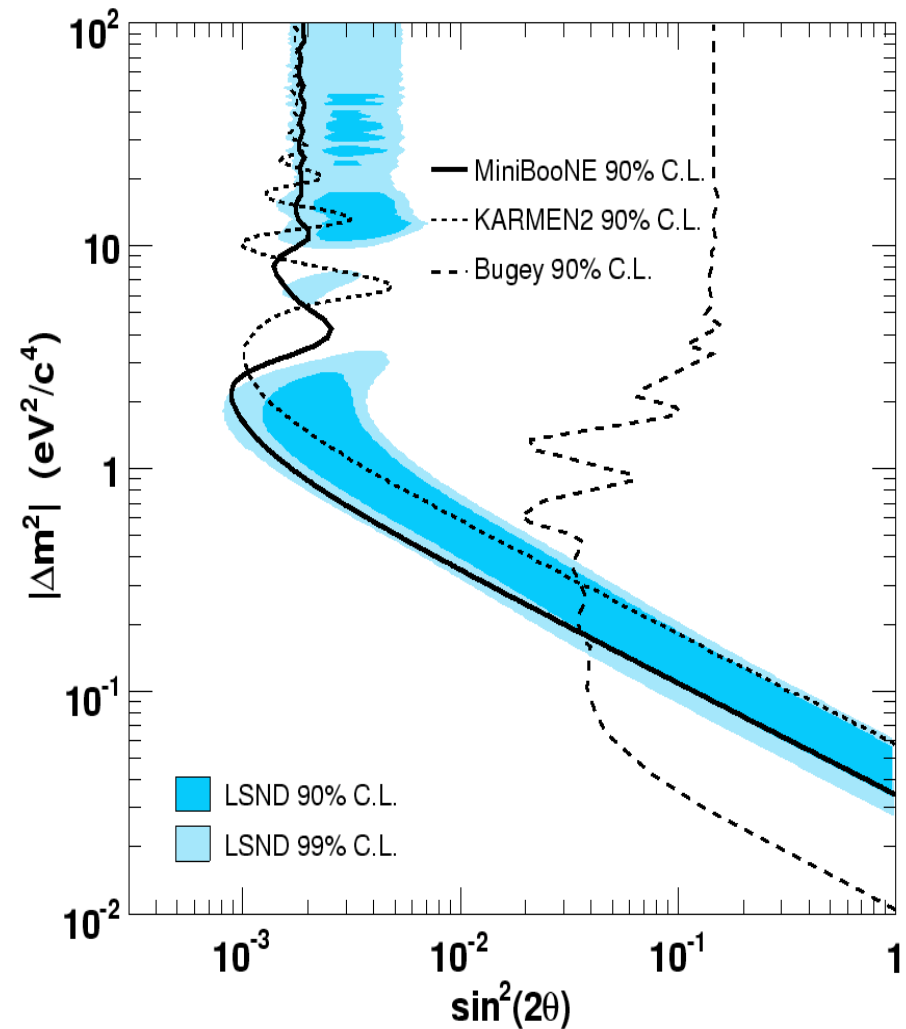
# LSND $\bar{\nu}_e$ appearance

## LSND experiment

Observation of  $3.8 \sigma$  excess  
of  $\bar{\nu}_e$  in  $\bar{\nu}_\mu$  beam

Karmen, Bugey and  
MiniBooNE exclude the  
LSND parameter space

If  $\bar{\nu}_e$  oscillate but  $\nu_e$  do  
not, then exotic physics is  
needed to explain this signal



# Sterile neutrinos

One explanation for the LSND oscillation signal is to add another “sterile” flavor of neutrino (or 2 or N) to the mixing matrix:

Adding 1 sterile neutrino is 3+1, adding N is 3+N

$$U_{\alpha i} = \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \\ \vdots \\ \nu_s \end{pmatrix} \begin{pmatrix} U_{e1} & U_{e2} & \cdots & U_{eN} \\ U_{\mu1} & U_{\mu2} & \cdots & U_{\mu N} \\ U_{\tau1} & U_{\tau2} & \cdots & U_{\tau N} \\ \vdots & \vdots & \ddots & \vdots \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \\ \vdots \\ \nu_N \end{pmatrix}$$

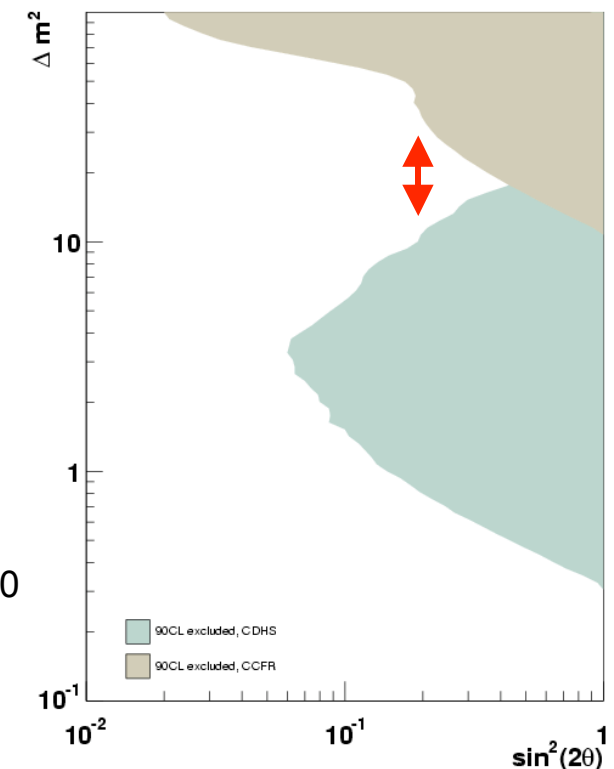
Disappearance expts (CDHS/CCFR/atmospheric)

disfavor 3+1 already

Maltoni, Schwetz, Valle, Phys.Lett.B518:252-260,2001. hep-ph/0107150

3+2 models have large mixing and prefer the region where experimental limits are weakest

G. Karagiorgi, V. Barger et al, Phys.Rev.D75:013011,2007. hep-ph/0609177



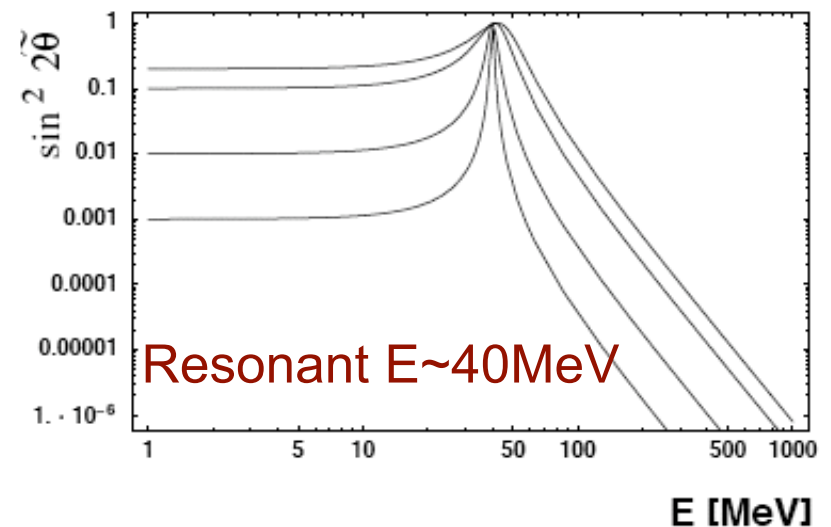
# Motivation for neutrino disappearance

The observation of  $\nu_\mu$  disappearance could imply:

- sterile neutrinos G. Karagiorgi et al, Phys.Rev.D75:013011,2007. hep-ph/0609177
- neutrino decay Palomares-Ruiz, Pascoli, Schwetz, JHEP 0509:048,2005. hep-ph/0505216
- **extra dimensions** Pas, Pakvasa, Weiler, Phys.Rev.D72:095017,2005. hep-ph/0504096

When the path-length increases for active neutrinos in the bulk relative to sterile neutrinos, **oscillations between sterile and active flavors are enhanced above a resonant energy**, and suppressed below

A resonance energy between 30-400MeV explains all data in a 3+1 model



The lack of  $\nu_\mu$  disappearance also can constrain these models

# Motivation for neutrino disappearance

The combination of  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance tests unitarity of the mixing matrix, and CPT

❖ If  $\bar{\nu}_\mu$  disappear, but  $\nu_\mu$  do not would signal CPT violation

$$\text{Prob}(\nu_\mu \rightarrow \nu_x) \neq \text{Prob}(\bar{\nu}_\mu \rightarrow \bar{\nu}_x)$$

❖ Sterile neutrino models (3+1 or 3+2) can be CPT violating

Barger, Marfatia, & Whisnant, Phys. Lett. B576 (2003) 303

❖ Introduction of a new light gauge boson

Nelson, Walsh Phys. Rev. D77 033001 (2008) hep-ph/0711.1363



# Motivation for neutrino disappearance

- The observation of  $\nu_\mu$  disappearance could imply new physics
- The lack of  $\nu_\mu$  disappearance constrains new physics models
- The combination of  $\nu_\mu$  and  $\bar{\nu}_\mu$  disappearance tests unitarity and CPT

Can MiniBooNE add to the current disappearance limits?

YES! with both neutrinos and antineutrinos

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- 6) Conclusion

# The MiniBooNE Collaboration

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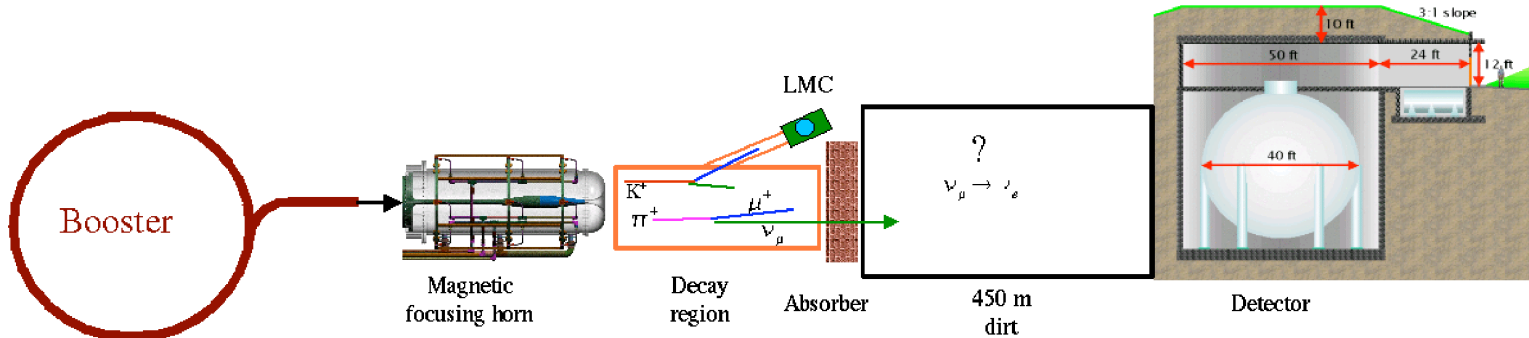
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# MiniBooNE Experiment

Short baseline ( $L \sim 500\text{m}$ ) designed to test LSND-like  $\nu_e$  appearance



8.9 GeV/c protons on Be produce mesons which decay to neutrinos

Booster:  $4 \times 10^{12}$  protons /  $1.6 \mu\text{s}$

pulse delivered at up to 5 Hz

$p + 1.7 \lambda \text{Be}$  target produces mesons

Magnetic horn focuses

mesons, pulsed at 174kA

Increases flux by  $\sim \times 6$

Decay region:  $\pi$ , K decay to neutrinos

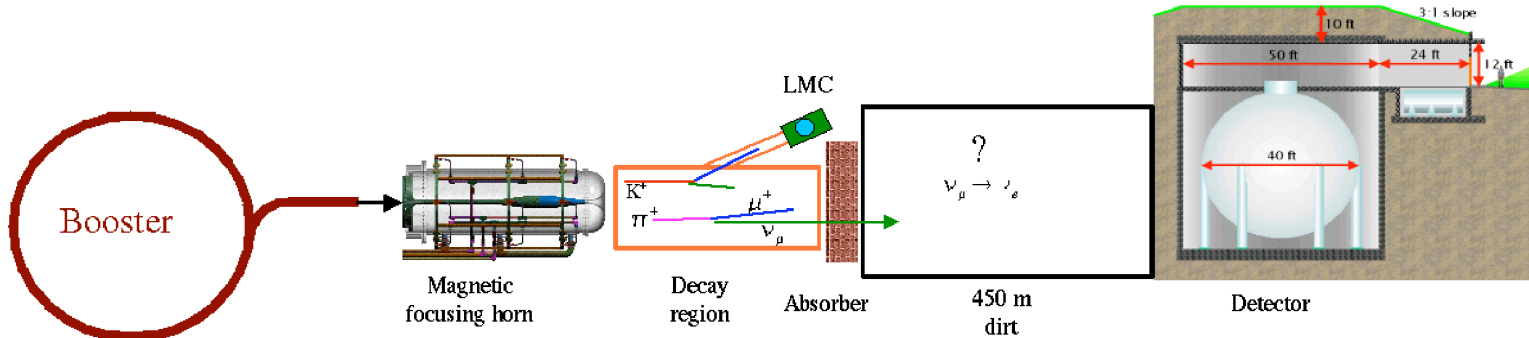
$\sim 450\text{m}$  of earth stops any

remaining particles

MiniBooNE detector

# MiniBooNE Experiment

Short baseline ( $L \sim 500\text{m}$ ) designed to test LSND-like  $\nu_e$  appearance



8.9 GeV/c protons on Be produce mesons which decay to neutrinos or antineutrinos

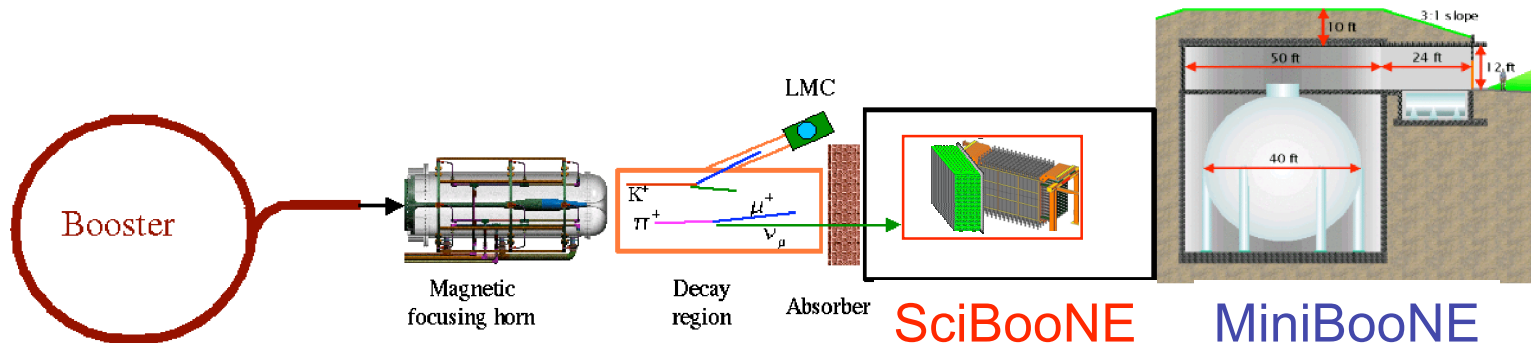
Changing the polarity of the horn focuses **positive** (**negative**) mesons and produces **a neutrino** (**antineutrino**) beam

Data sets shown today are:

**5.579e20POT neutrino mode** (190,454 events)

**3.386e20POT antineutrino mode** (27,053 events)

# Addition of SciBooNE Experiment



In May 2007, the SciBooNE detectors started taking data at 100m  
In August 2008, after two joint neutrino and antineutrino runs  
with MiniBooNE, SciBooNE was decommissioned

In MiniBooNE:

~1  $\nu$  per  $1e15$  POT

~ 0.2  $\bar{\nu}$  per  $1e15$  POT

In SciBooNE: ~5x closer, ~50x smaller

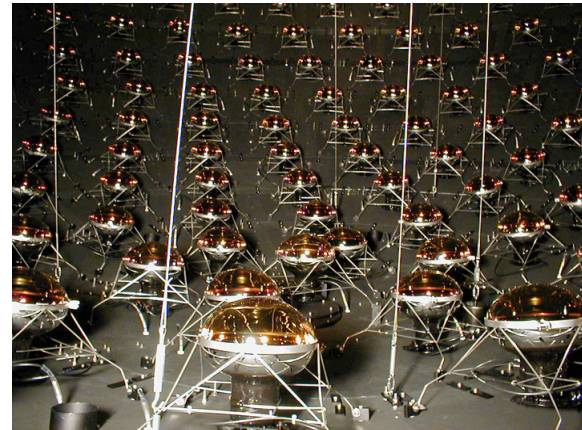
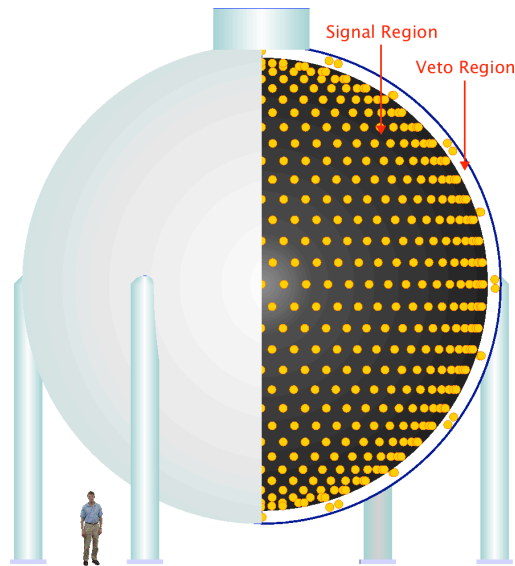
~0.5  $\nu$  per  $1e15$  POT

~ 0.1  $\bar{\nu}$  per  $1e15$  POT

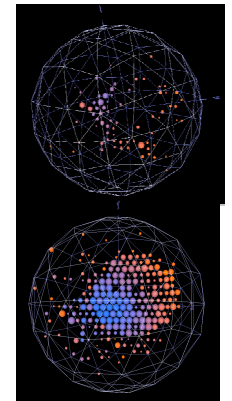
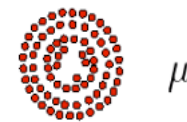
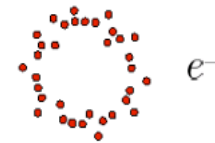
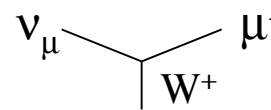
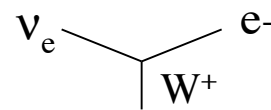
# MiniBooNE Detector

The MiniBooNE detector is a ~1kton mineral oil Cherenkov detector  
12 m diameter, 1280 inner PMTs, 240 outer 'veto' PMTs

MiniBooNE Detector

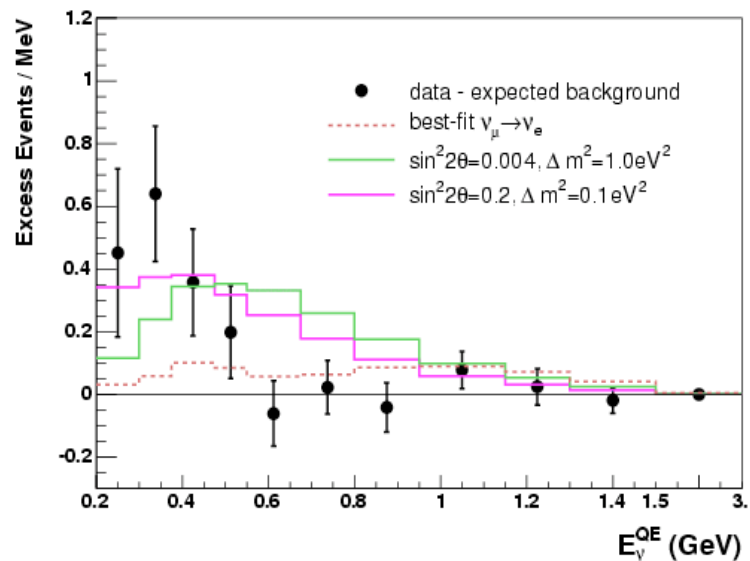
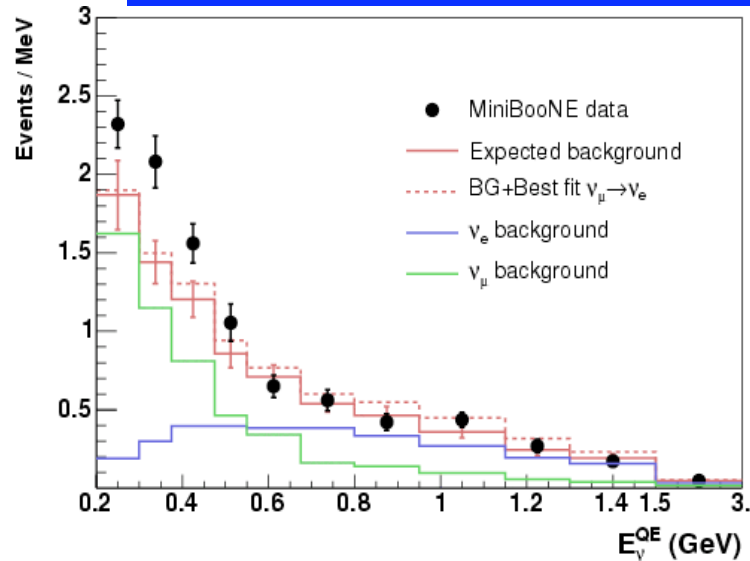


Use hit topology and timing to  
determine electron-like or muon-like  
Cherenkov rings and corresponding  
charged current neutrino interactions





# MiniBooNE $\nu_e$ appearance results



- $\nu_e$  sample is consistent with expectation  $>475 \text{ MeV}$  ( $0.6 \sigma$  excess)

- $3.0 \sigma$  excess at low energy ( $200\text{--}475 \text{ MeV}$ )

Initial observation confirmed with later work as presented this August; PRL forthcoming

Excess cannot be described based on a simple 2  $\nu$  mixing hypothesis

- This result assumes no  $\nu_{\mu}$  disappearance



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# $\nu_\mu$ disappearance analysis plan

To do a  $\nu_\mu$  disappearance analysis with one detector, we need:

Event selection

+

Prediction with systematic errors  
flux, cross section, detector effects

+

Disappearance fit machinery

# $\nu_\mu$ disappearance sample

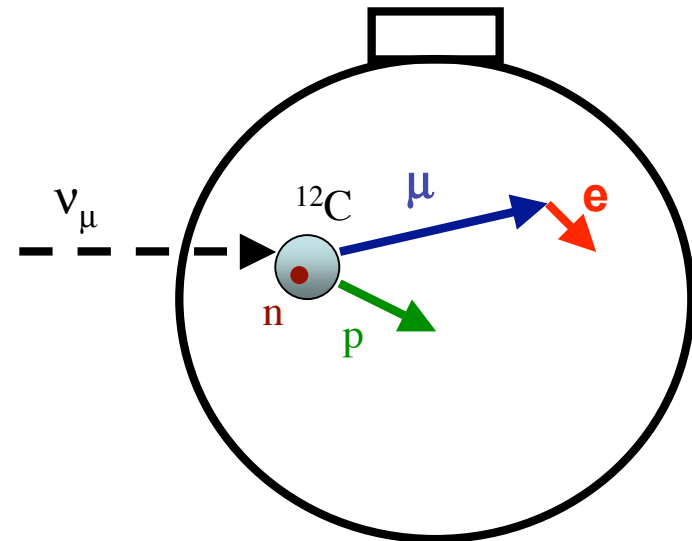
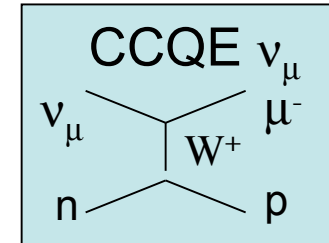
- Use Charged Current Quasi elastic events (CCQE)  $\nu_\mu$  events

- Selecting on muon selects  $\nu_\mu$
- With just muon's energy, angle, can reconstruct neutrino energy

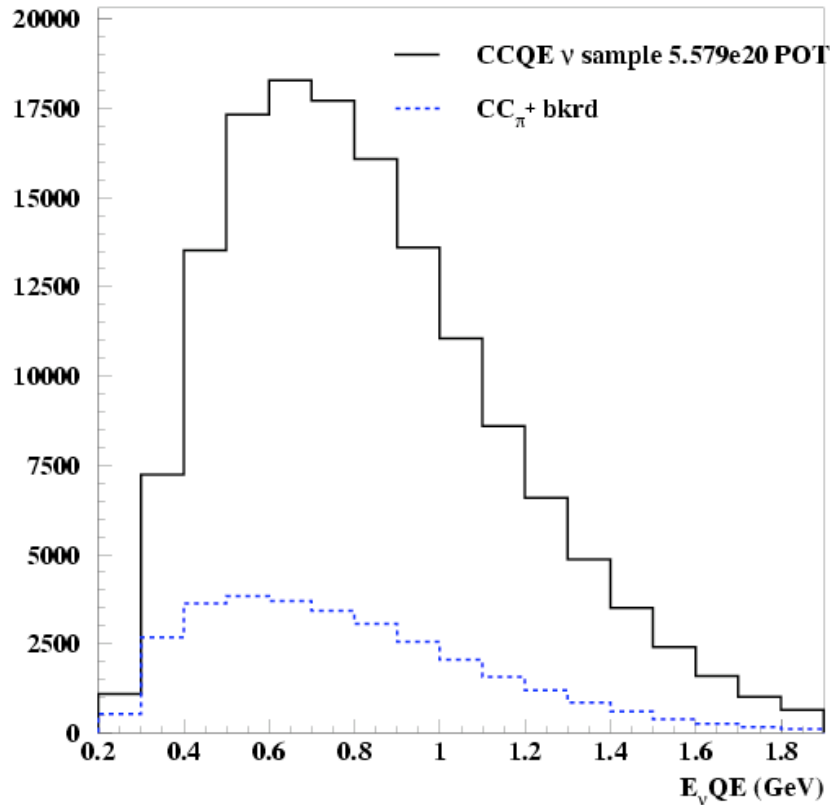
$$E_\nu(QE) = \frac{m_n E_\mu - \frac{1}{2} m_\mu^2}{|p_\mu| \cos \theta_\mu + m_n - E_\mu}$$

## Tag single muon events and their decay electron

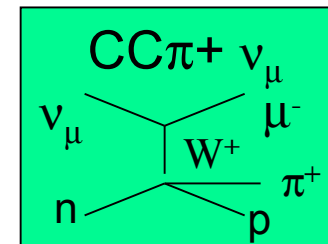
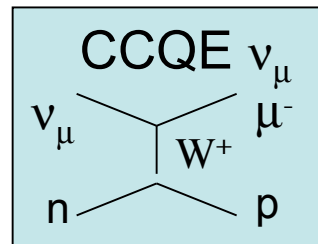
- 2 subevents ( $\mu$ , then  $e$ ) with minimal veto activity in both
- muon-like track, 2nd event below decay electron energy endpoint
- both events within fiducial volume



# CCQE $\nu_\mu$ selection



■ Impressive neutrino sample: ~200k events, 74% CCQE purity



- Background is  $CC\pi^+$  where the pion is absorbed in the nucleus or detector
  - All events can oscillate, but misreconstruction of  $CC\pi^+$  as CCQE events mean  $CC\pi^+$  are shifted to low  $E_{\nu QE}$
- Pure neutrino sample, only 1.4% antineutrino content

# $\nu_\mu$ disappearance analysis plan

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flux, cross section, detector effects

+

Disappearance fit machinery

# Flux prediction

Neutrino beamline is modeled in Geant4 hep-ex/0806.1449

p + Be target → meson production → focusing → decay → neutrinos

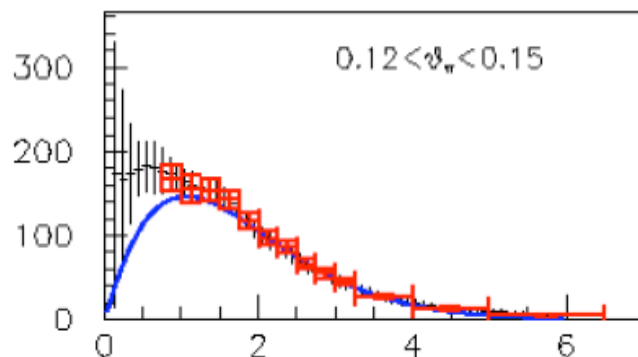
Included as systematic error:

1. Beam optics and targeting efficiency
2. p+Be elastic and inelastic cross sections
3. Production of mesons ( $\pi^{+/-}$ ,  $K^{+/-}$ ) from pBe interactions
4. Horn magnetic field

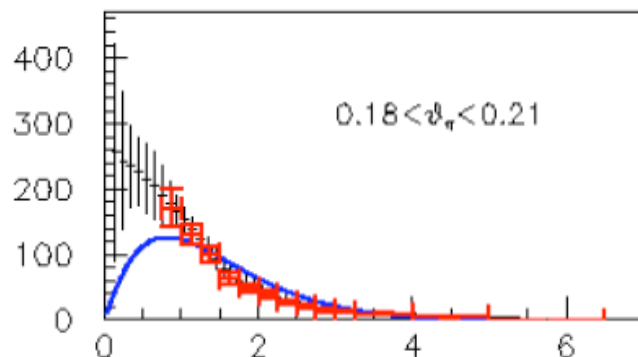
Largest sources of error are meson production and horn magnetic field

# Meson Production Uncertainties

$d\sigma/dp d\Omega$  (mb c/[GeV sr])



$p_\pi$  (GeV)



HARP data with errors in  $\theta_\pi$  bins  
MiniBooNE flux parameterization

The HARP experiment measured  
 $p+\text{Be} \rightarrow \pi^+/\pi^-$  (hep-ex/0702024)

Use the HARP data and errors to  
produce different fluxes consistent  
with HARP

Propagate the new fluxes through to  
the neutrino spectrum and look at the  
effect on the CCQE  $\nu_\mu$  sample

88% of the CCQE  $\nu_\mu$  sample is within  
HARP's coverage; 99% is contained  
within HARP and  $\theta_\pi > 0.210$

# Cross section model and the disappearance result

For  $\nu_e$  appearance result, we tuned the cross section model

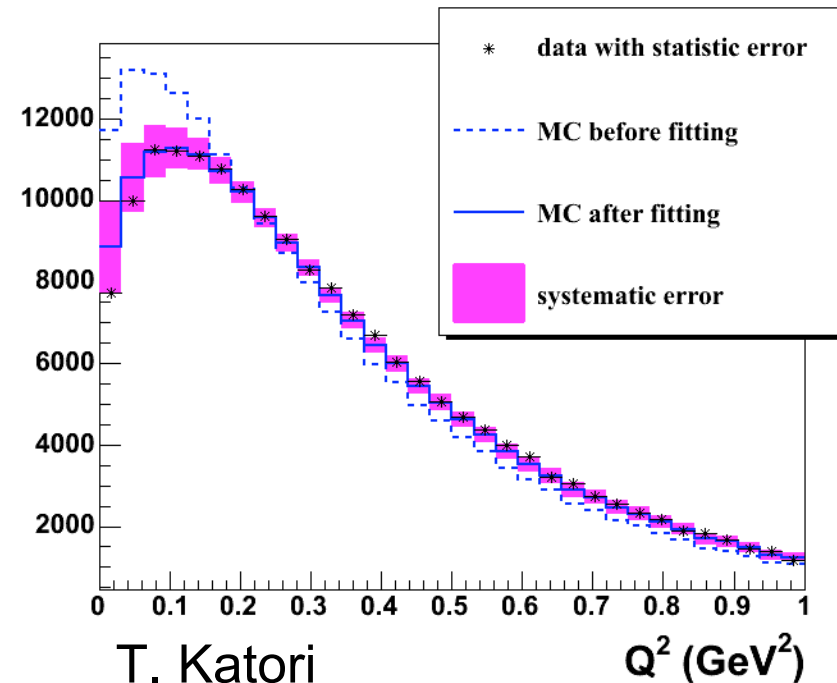
Shape only fit in  $Q^2$  using the CCQE  $\nu_\mu$  sample favored a higher axial form factor ( $M_A$ ) and a new nuclear effect parameter,  $K$ , was introduced to model Pauli suppression or other effects at low  $Q^2$

Phys. Rev. Lett. 100, 032301 (2008).

$$M_A = 1.23 \pm 0.20 \text{ GeV}$$

$$K = 1.019 \pm 0.011$$

$$Q^2 = -m_\mu^2 + 2E_\nu(E_\mu - p_\mu \cos \theta_\mu)$$





## Cross section model and the disappearance result

For  $\nu_\mu$  disappearance, we undo the tuning and set the uncertainties to cover the excursion in the world data and our own

World's data on deuterium:  $M_A = 1.014 \pm 0.014$  GeV

Bodek et al J.Phys.Conf.Ser.110:082004,2008. hep-ex/0709.3538

K2K CCQE  $\sigma$  on Carbon:  $M_A = 1.14 \pm 0.11$  GeV

F. Sanchez, NuInt07

K2K CCQE  $\sigma$  on Oxygen  $M_A = 1.20 \pm 0.12$  GeV

R. Gran et al., PRD74, 052002 (2006)

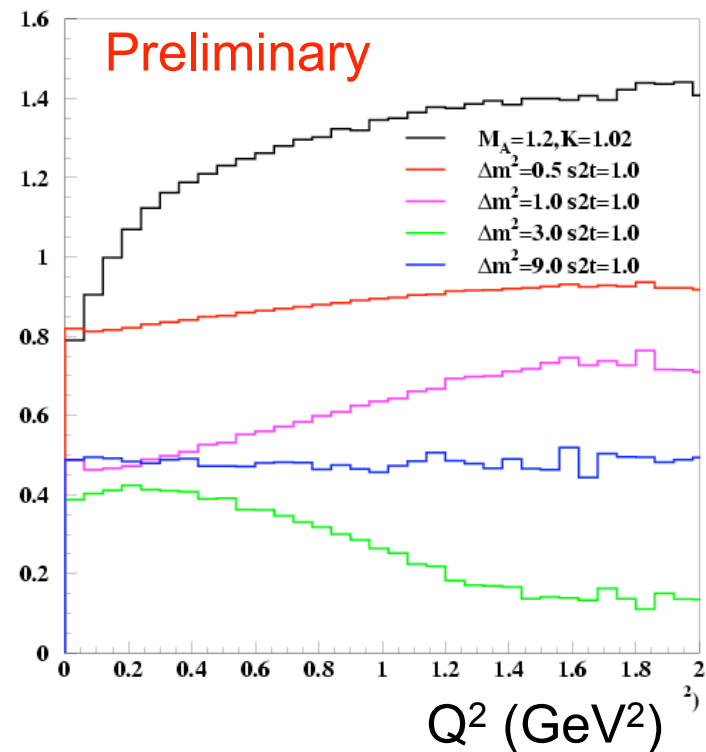
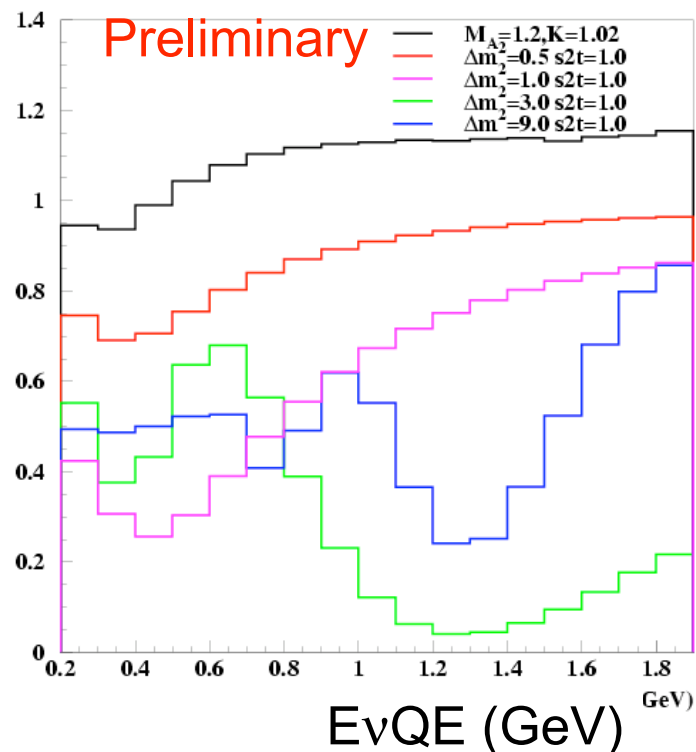
Using:  $M_A = 1.0 \pm 0.23$  GeV,  $K = 1.000 \pm 0.0220$

The cross section uncertainties also include uncertainties on the  $CC\pi^+$  cross section and pion charge exchange and absorption in the nucleus

## Can the cross section model mask disappearance?

$(M_A=1.2 \text{ GeV}, K=1.02) / (M_A=1.0 \text{ GeV}, K=1.0)$  induces a shape change similar to  $\Delta m^2=0.5 \text{ eV}^2$  in EvQE

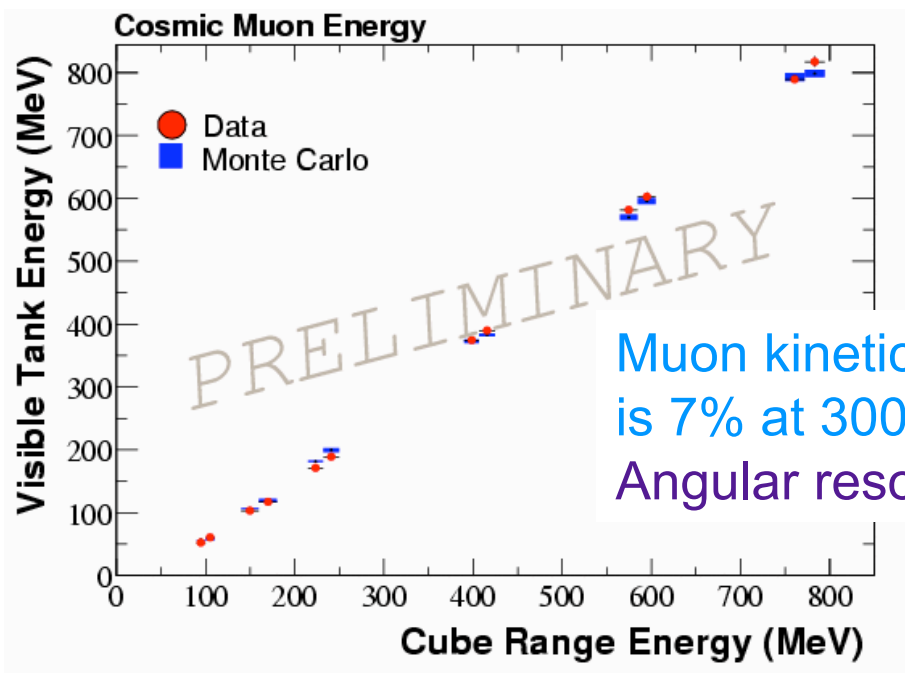
But in  $Q^2$ , oscillations vanish while the effect of the cross sections is stronger



# Detector uncertainties

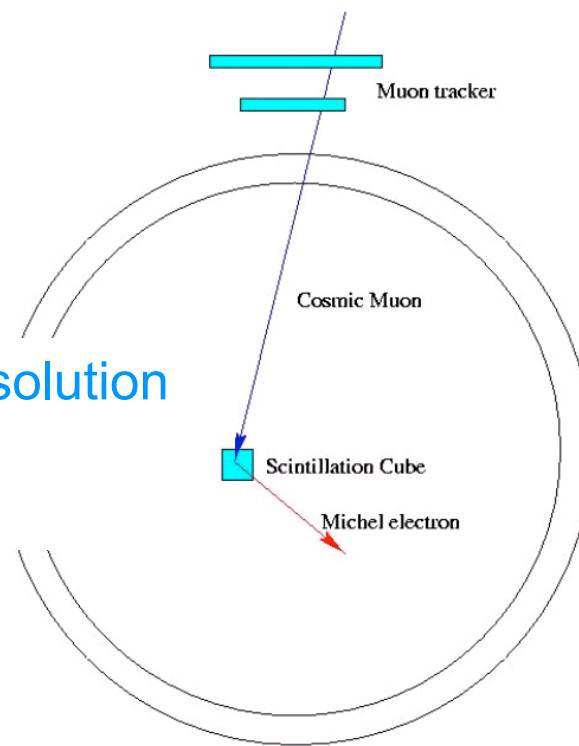
Muon hodoscope tracked incoming (10kHz) cosmic ray muons entering detector

Events which stopped in scintillation cubes provided known distance with which to calibrate muon energy in oil



Muon kinetic energy resolution is 7% at 300MeV

Angular resolution is  $5^\circ$



# Systematic error summary

Source of error	Total fractional error (%) (counting experiment)
pBe $\rightarrow$ $\pi^+$ production (flux)	4.0
beamline and horn model	4.3
cross sections	18.6
detector model	4.0
total	19.9

Data = 190,454 events

MC (MA,K=1.0) = 145,085 +/- 20%

- The more one under predicts the data, the stronger the sensitivity to  $\nu_\mu$  disappearance becomes
- We under predict the data normalization by  $1.5 \sigma$
- In order to be conservative, we choose to perform a shape only disappearance fit
- Normalization information will be included with SciBooNE

# $\nu_\mu$ disappearance analysis plan

To do a  $\nu_\mu$  disappearance analysis with one detector, we need:

Event selection

+

Prediction with systematic errors  
flux, cross section, detector effects

+

Disappearance fit machinery

# Shape-only disappearance fit

Use Shape only Pearson's  $\chi^2$ :

For each point in oscillation space compare the prediction,  $p_i(\Delta m^2, \sin^2 \theta)$ , to the data,  $d_i$ , and sum over bins  $i$  and  $j$

$$\chi^2 = \sum (d_i - X p_i) M_{ij}^{-1} (d_j - X p_j)$$

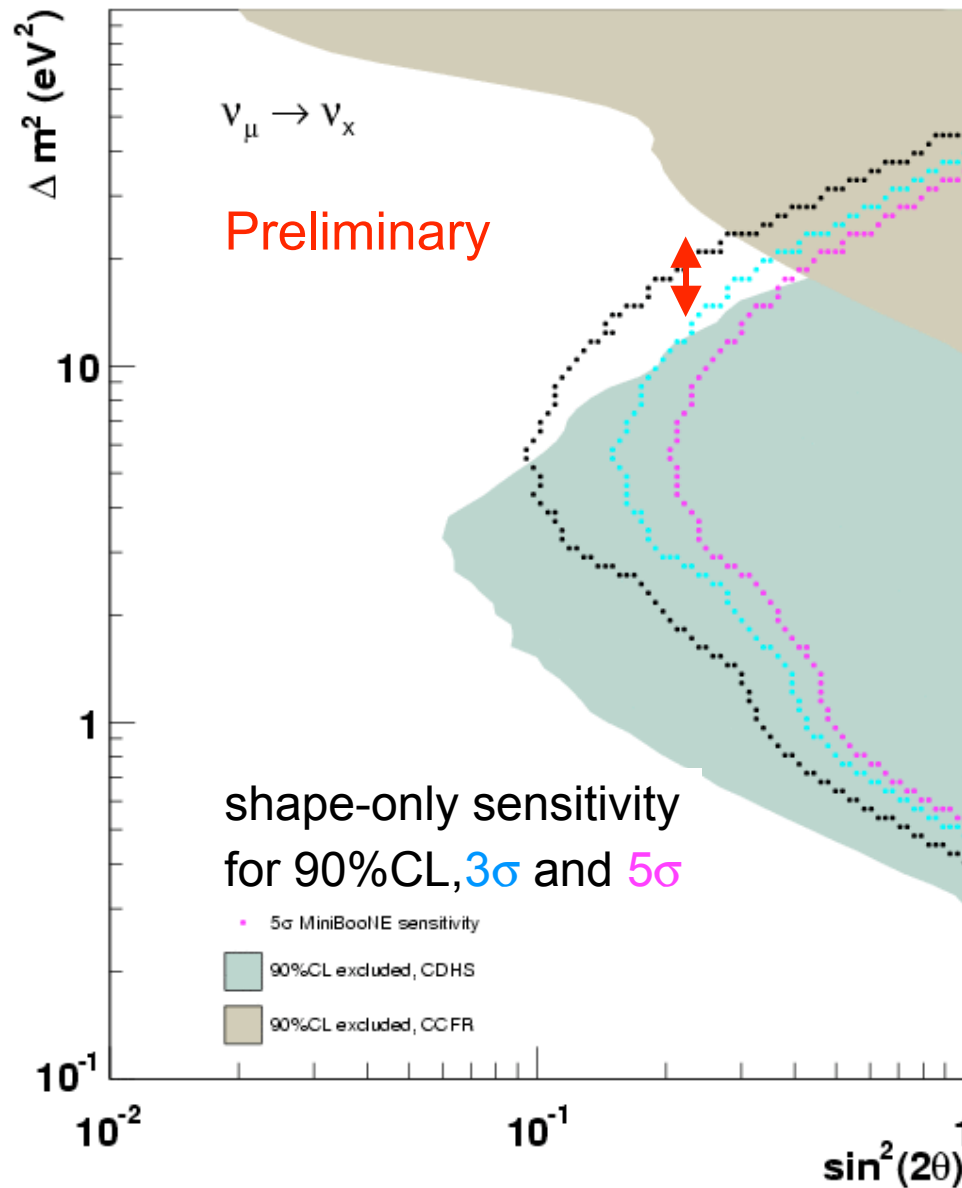
- $M_{ij}$  is shape only (variations conserve events across all bins)
- $X(\Delta m^2, \sin^2 \theta)$  renormalizes  $p_i$  to the total data events,

$$X(\Delta m^2, \sin^2 2\theta) = \frac{\sum d_i}{\sum p_i}$$

For  $\Delta m^2, \sin^2 \theta$  points where  $\chi^2 > \chi^2(\text{CL})$ , draw that CL curve

For 16 bins,  $\chi^2(90\% \text{ CL}) = 23.5$

# Sensitivity



The sensitivity is a fit to fake data which exactly agrees with prediction but all statistical and systematic uncertainties are included

A shape-only, single detector measurement is sensitive to  $\nu_\mu$  disappearance in the particular region favored by **3+2 models**

# Cross check: Frequentist $\Delta\chi^2$

Comparison between data ( $d_i$ ) and prediction ( $p_i$ ) relative to best fit across all  $\Delta m^2, \sin^2\theta$  points

$$\chi^2 = \sum (d_i - Xp_i) M_{ij}^{-1} (d_j - Xp_j)$$

For each point, create 50 “fake experiments” using fluctuations consistent with the errors and calculate  $\Delta\chi^2(\Delta m^2, \sin^2\theta, CL)$

$$\Delta\chi^2(\Delta m^2, \sin^2 2\theta) = \chi^2(true = \Delta m^2, \sin^2 2\theta) - \chi^2(best)$$

For fit to real data, use  $\Delta\chi^2(\Delta m^2, \sin^2\theta, CL)$  to generate CL curves

Fit data at each point as if it corresponds to that true point, calculate  $\Delta\chi^2$

if  $\Delta\chi^2 > \Delta\chi^2(\Delta m^2, \sin^2\theta, CL)$  for a given CL, draw curve

Procedure can be done with shape-only fits like Pearson's  $\chi^2$

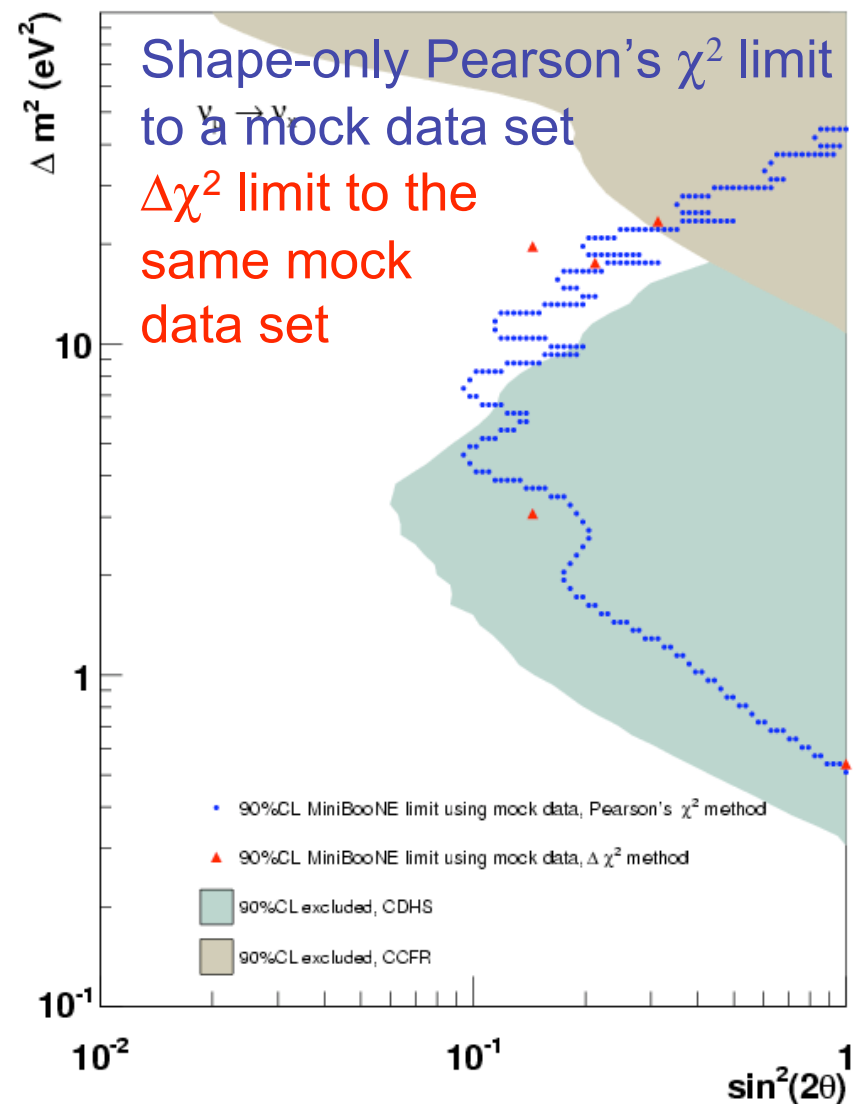
Renormalize  $p_i$  at each point, matrix is shape only



# Cross check: Frequentist $\Delta\chi^2$

Frequentist  $\Delta\chi^2$  gives better sensitivity by mapping out distorted  $\Delta\chi^2$  surface but is computing intensive

- $\Delta\chi^2$  ranges from  $\sim 4$  degrees of freedom (dof) at low  $\sin^2\theta$  to 1dof at high  $\sin^2\theta$
- Approximately 1 hour of computing for each  $\Delta\chi^2$  point shown, as compared to the  $\sim 1$  minute needed for the Pearson's  $\chi^2$  limit

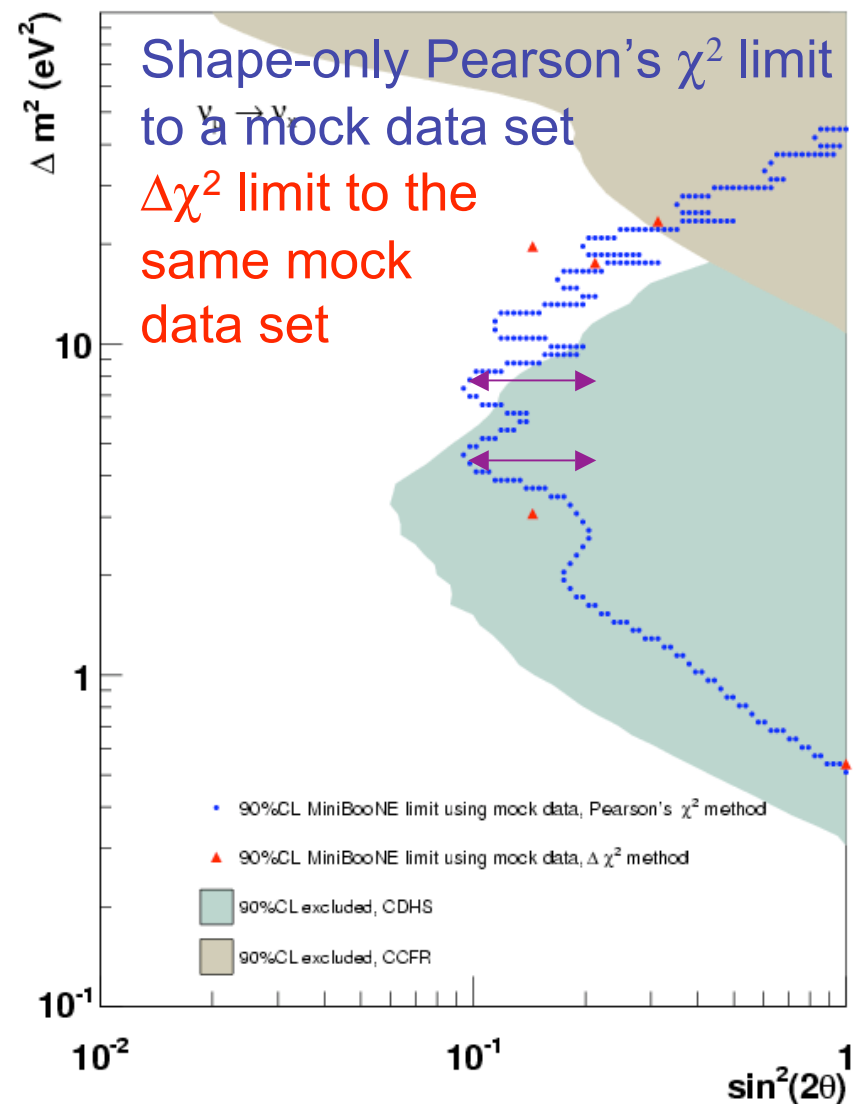
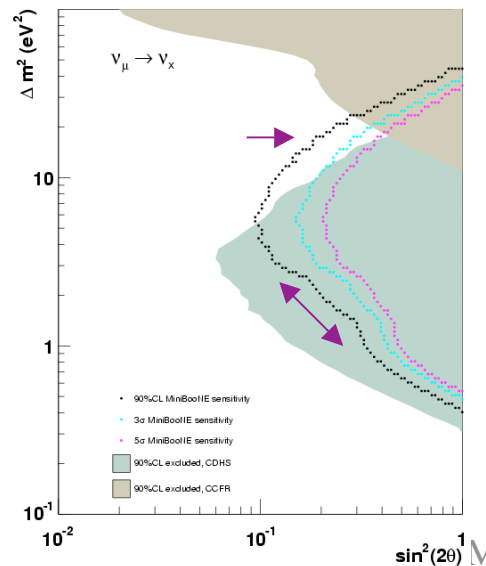


# Why does the limit look weird?

For all fits, the sensitivity curve can shift rapidly across  $\sin^2\theta$

We have been calling them “wiggles”

Wiggles are less pronounced in the sensitivity, but are present for any fake or real data fit



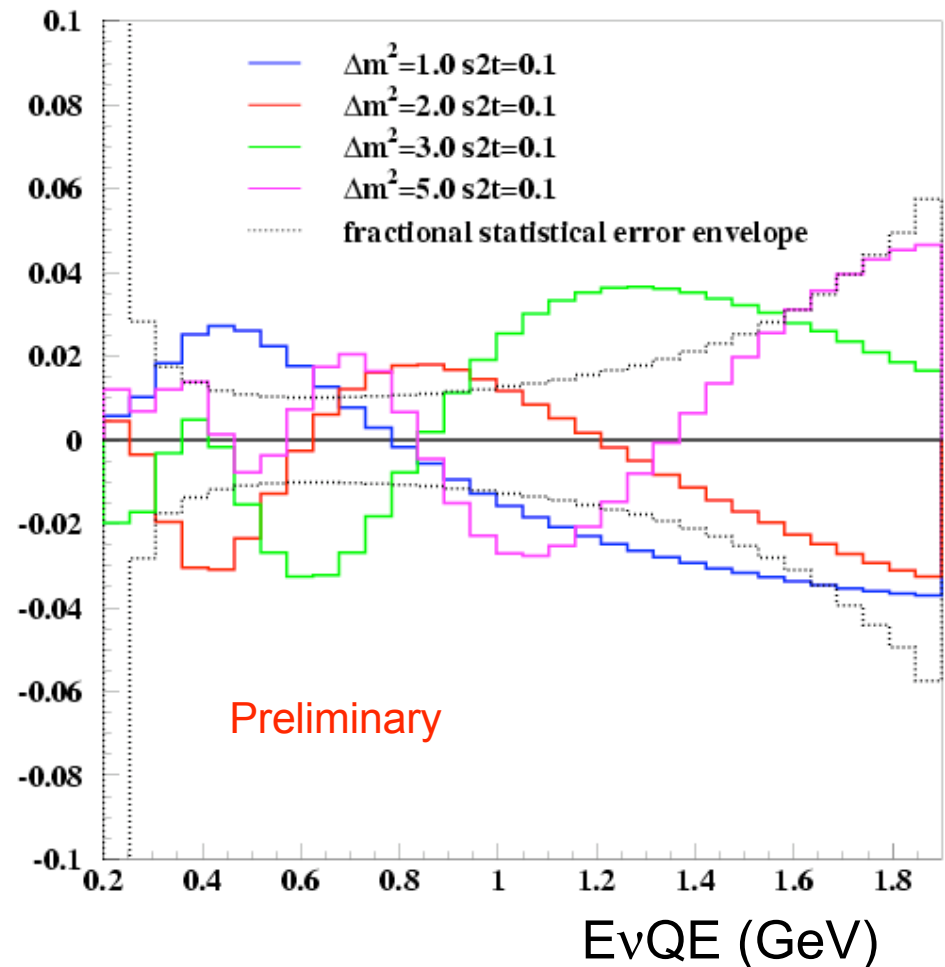
# What are the wiggles?

For a fixed  $\sin^2\theta$ ,  $\Delta m^2$  close in value do not have similar behavior in EvQE

For  $\sin^2\theta=0.1$ , if we compare the shape of  $\Delta m^2=2 \text{ eV}^2$  to  $\Delta m^2=3 \text{ eV}^2$  we see that the  $\chi^2(\Delta m^2=2 \text{ eV}^2) < \chi^2(\Delta m^2=3 \text{ eV}^2)$

The  $\chi^2$  changes with  $\Delta m^2$ ;  
a flat cut on  $\chi^2$  creates wiggles

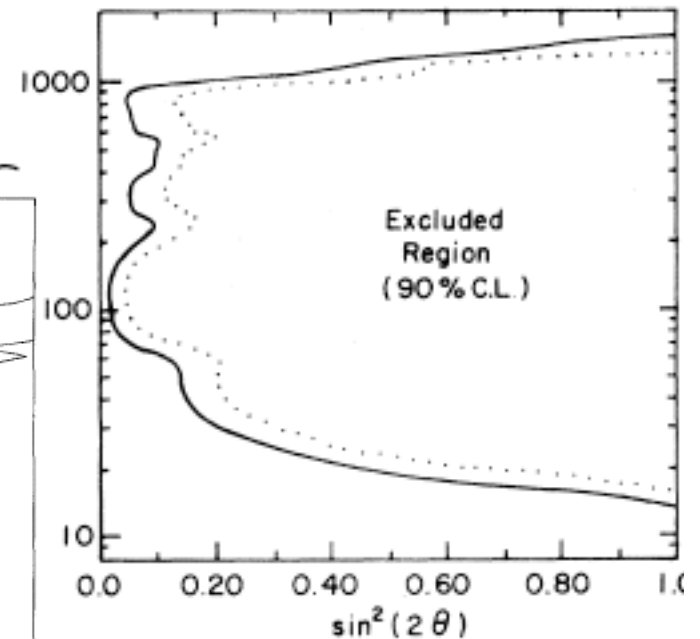
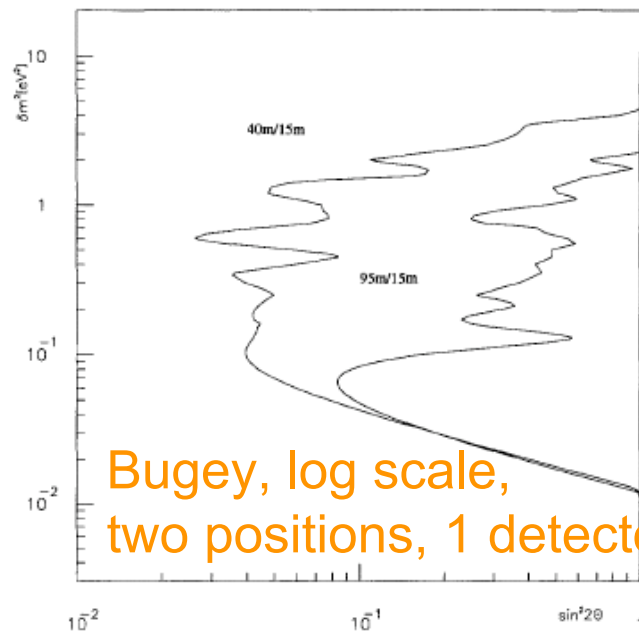
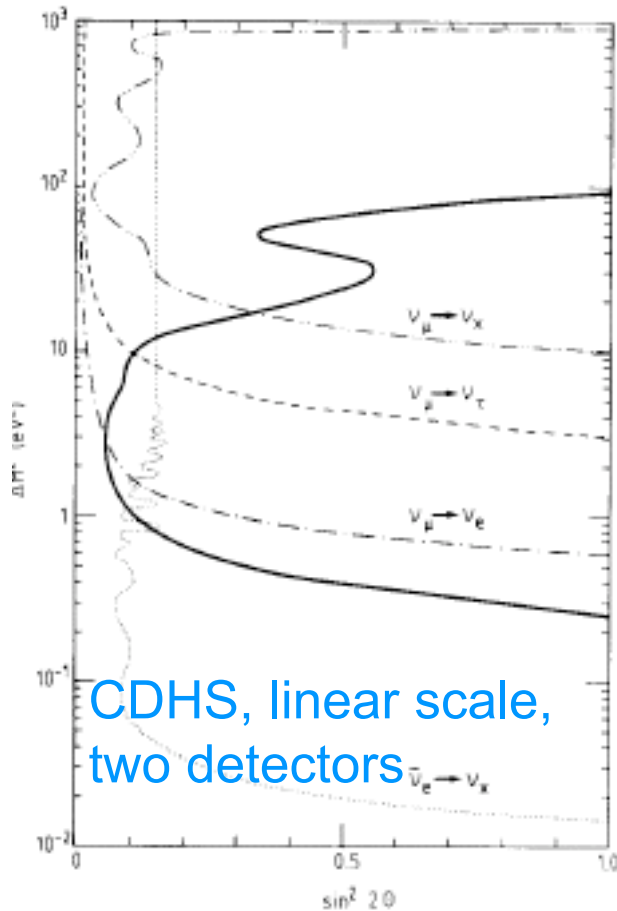
This problem is exacerbated for data fluctuations and can occur for any error envelope



# What are the wiggles?

This effect shows up in previous disappearance results even when there is a second detector

A second detector makes it harder to match L/E across all L, E but anytime it can, the  $\chi^2$  will be lower than nearby  $\Delta m^2$



# $\nu_\mu$ disappearance analysis plan

To do a  $\nu_\mu$  disappearance analysis with one detector, we need:

Event selection

+

Prediction with systematic errors  
flux, cross section, detector effects

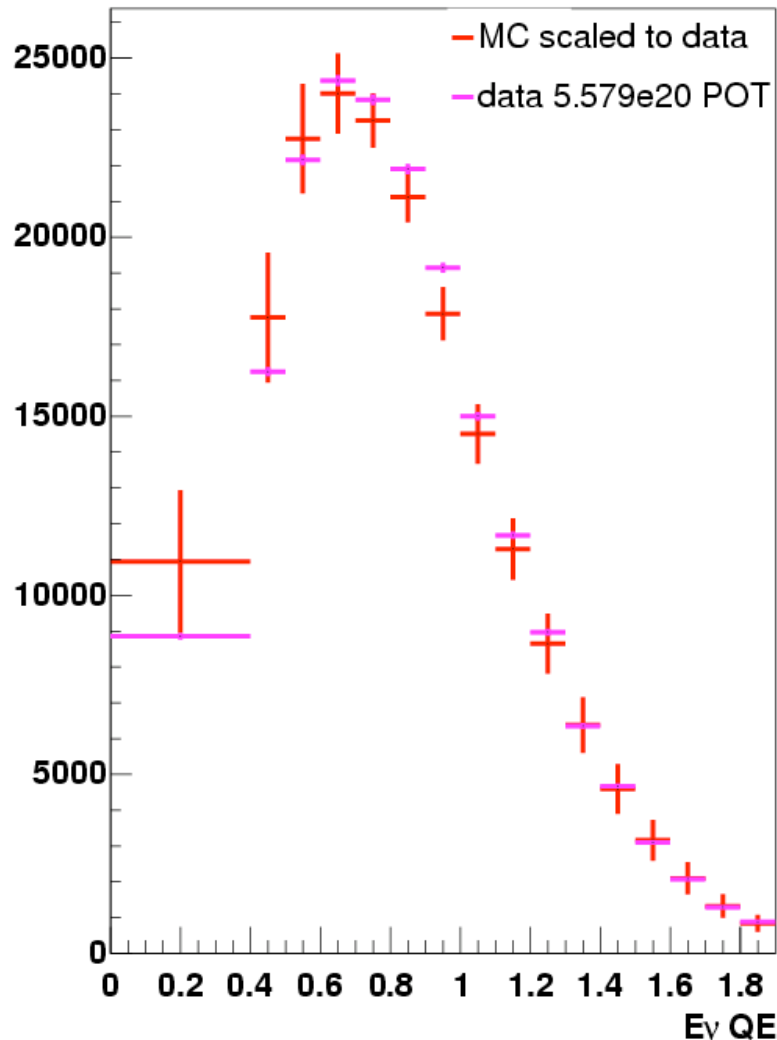
+

Disappearance fit machinery

=

Results!

# Data and null oscillation prediction



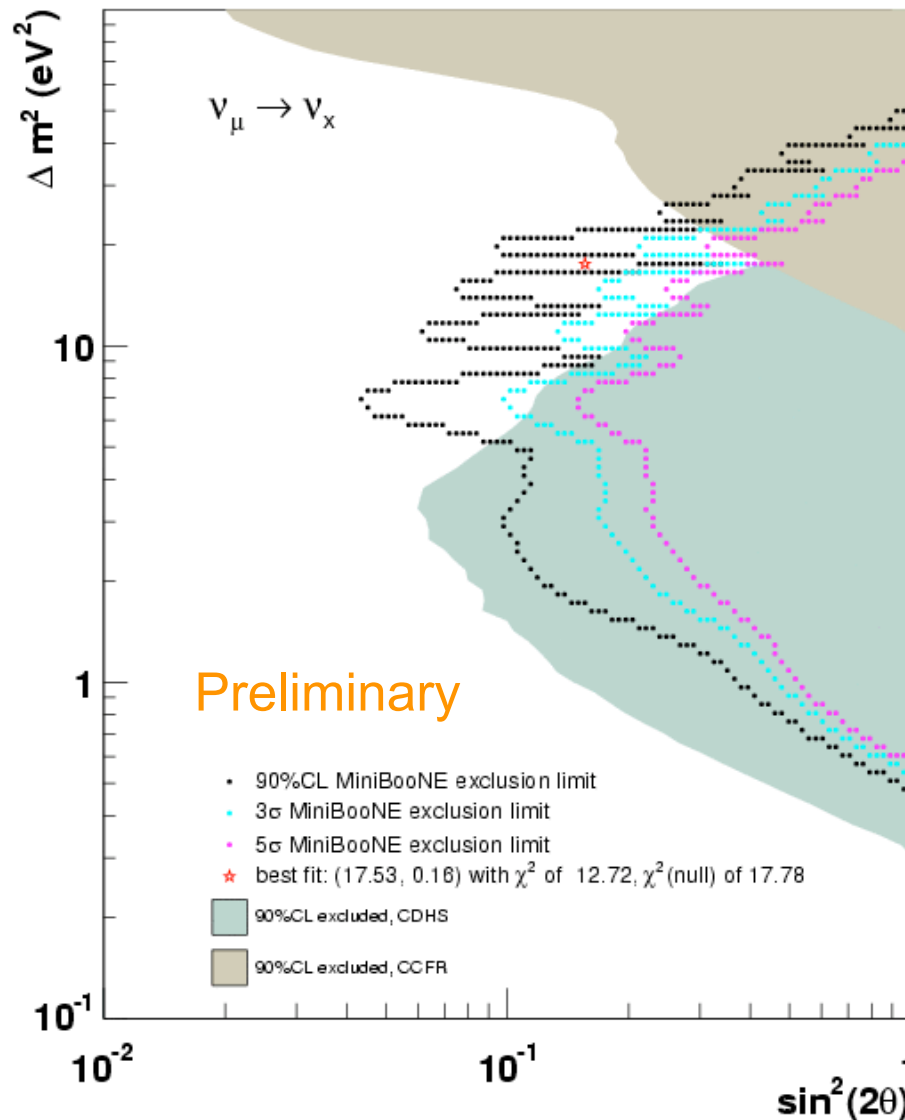
Data (5.579e20 POT, statistical errors shown) with null oscillation prediction (normalized to total data) vs  $E_{\nu}QE$

Errors shown are diagonal elements of the shape-only error matrix

$\chi^2(\text{null}) = 17.78$  (34% for 16 bins)

Systematics dominate:  
 $\chi^2(\text{null, statistics only}) = 665$

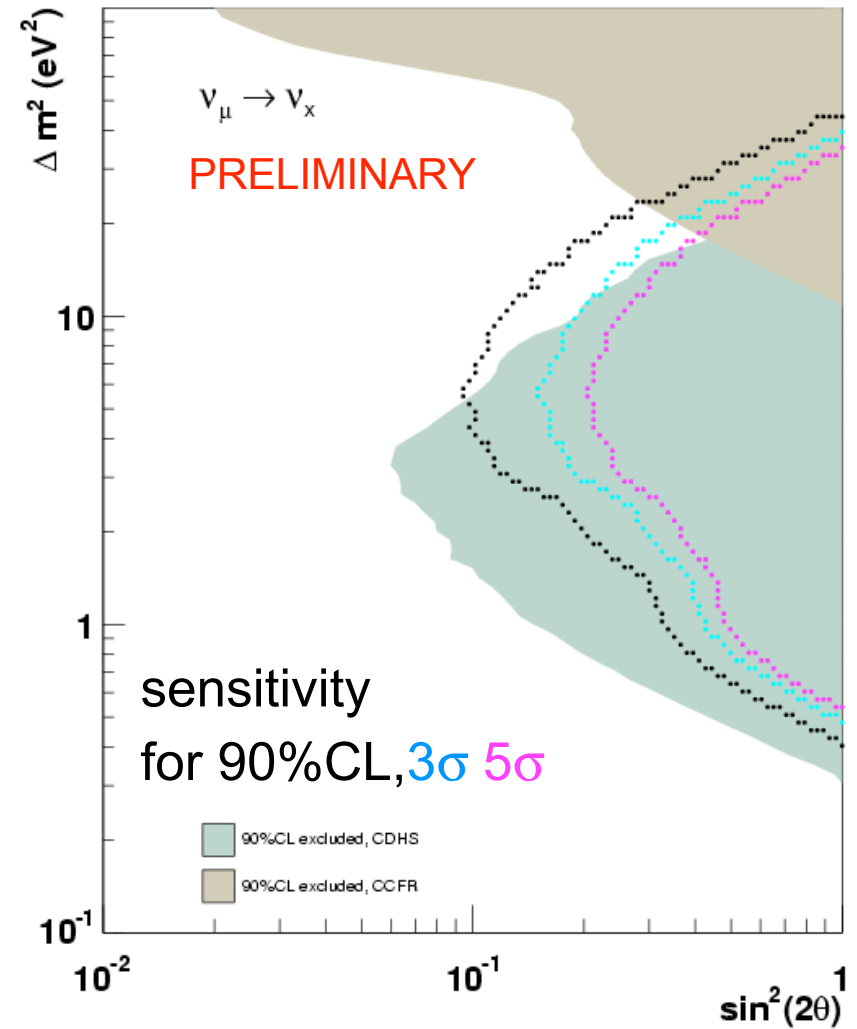
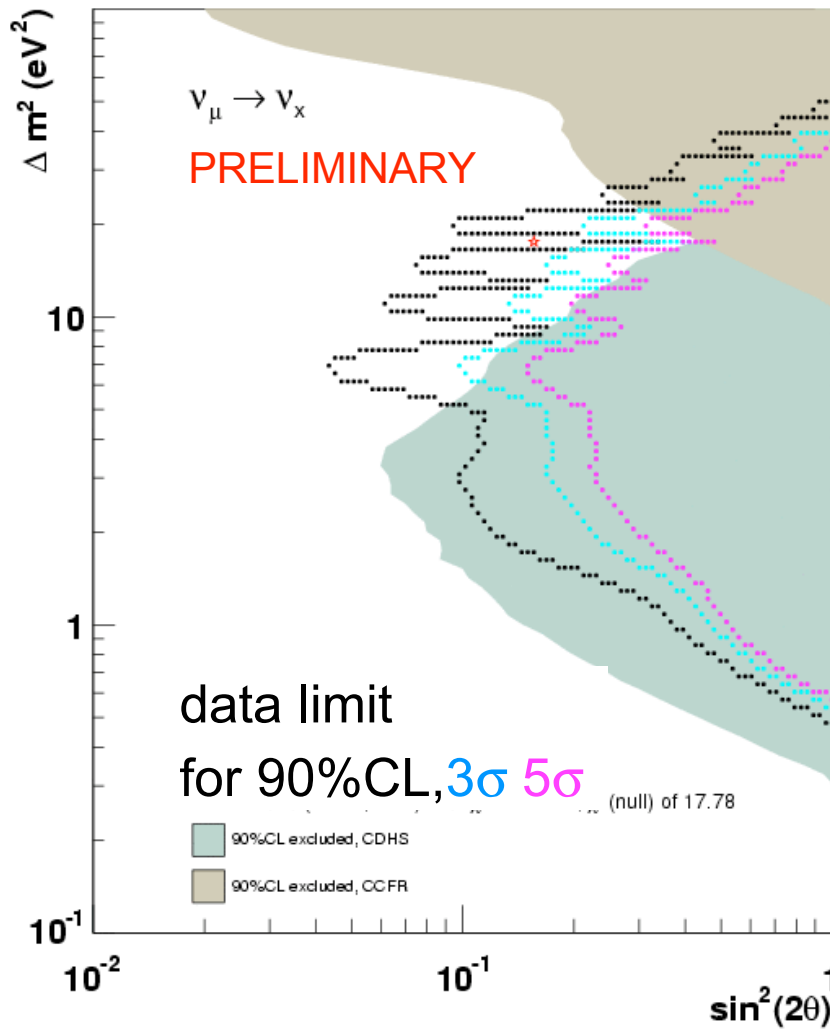
# Neutrino disappearance limit



5.579E20 POT data set limit  
for 90%CL,  $3\sigma$  and  $5\sigma$   
 $\chi^2$  (null) = 17.78 (34%, 16 bins)  
 $\chi^2$  (min) = 12.72 (69%, 16 bins)  
at  $\Delta m^2 = 17.5 \text{ eV}^2$ ,  $\sin^2 \theta = 0.16$

MiniBooNE observes  
no neutrino disappearance

# Neutrino disappearance limit



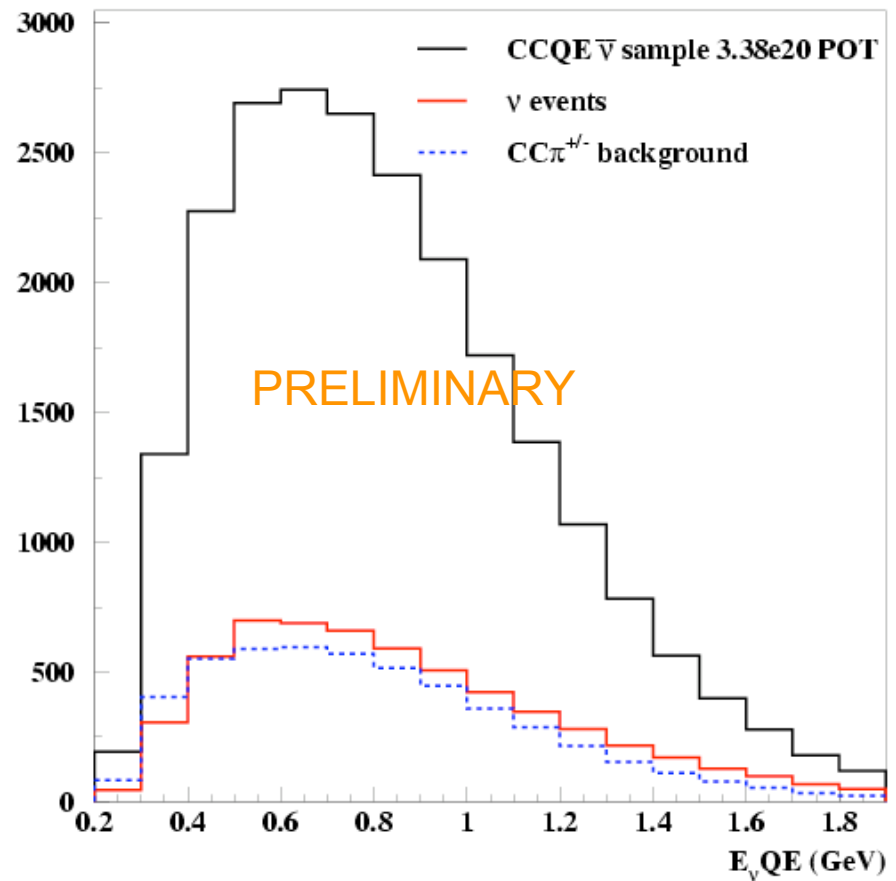


# Overview

- 1) Neutrino oscillation
- 2) MiniBooNE experiment
- 3) MiniBooNE-only neutrino disappearance analysis
- 4) Antineutrino disappearance analysis
- 5) Improvements to disappearance analysis
- 6) Conclusion

# Antineutrino CCQE sample

- Ability to change polarity of horn allows us to focus negative mesons and produce an antineutrino beam
- Apply same CCQE selection cuts, same error analysis, same fit machinery
- Main difference:  
Substantial neutrino events in the antineutrino sample (25%)

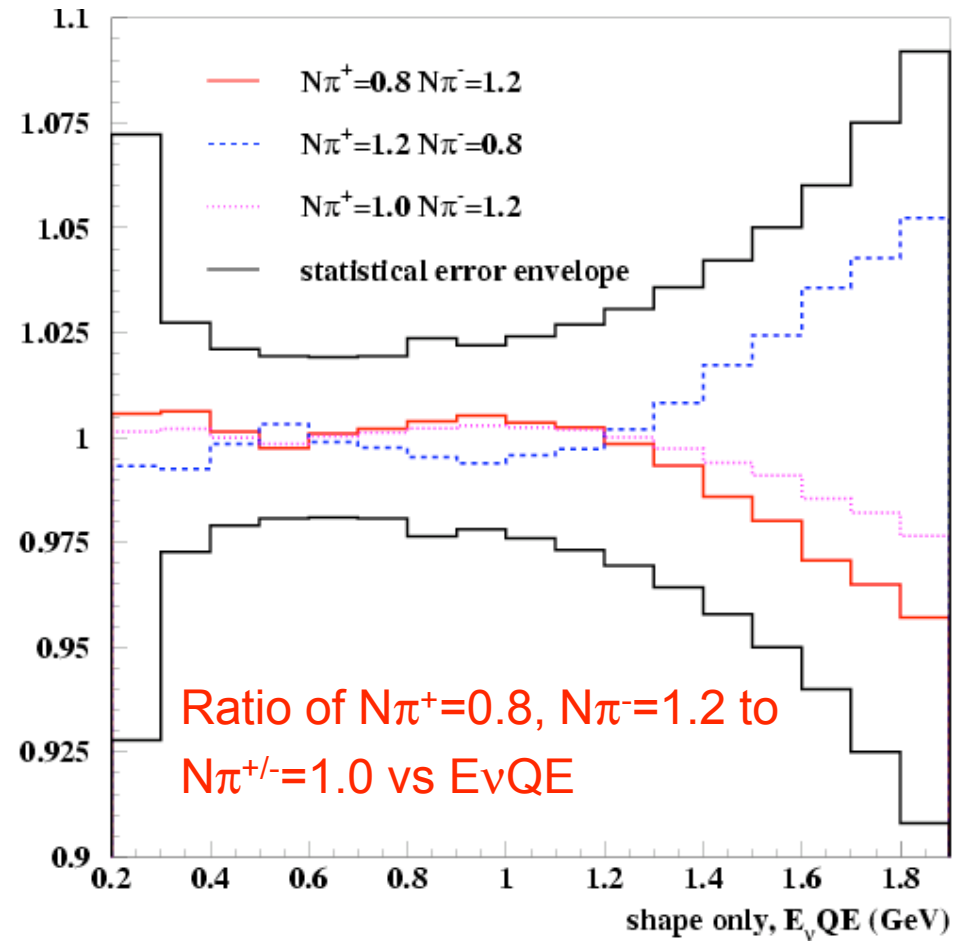


# Neutrinos in antineutrino sample

Is there a shape difference between the neutrino background and the antineutrino signal?

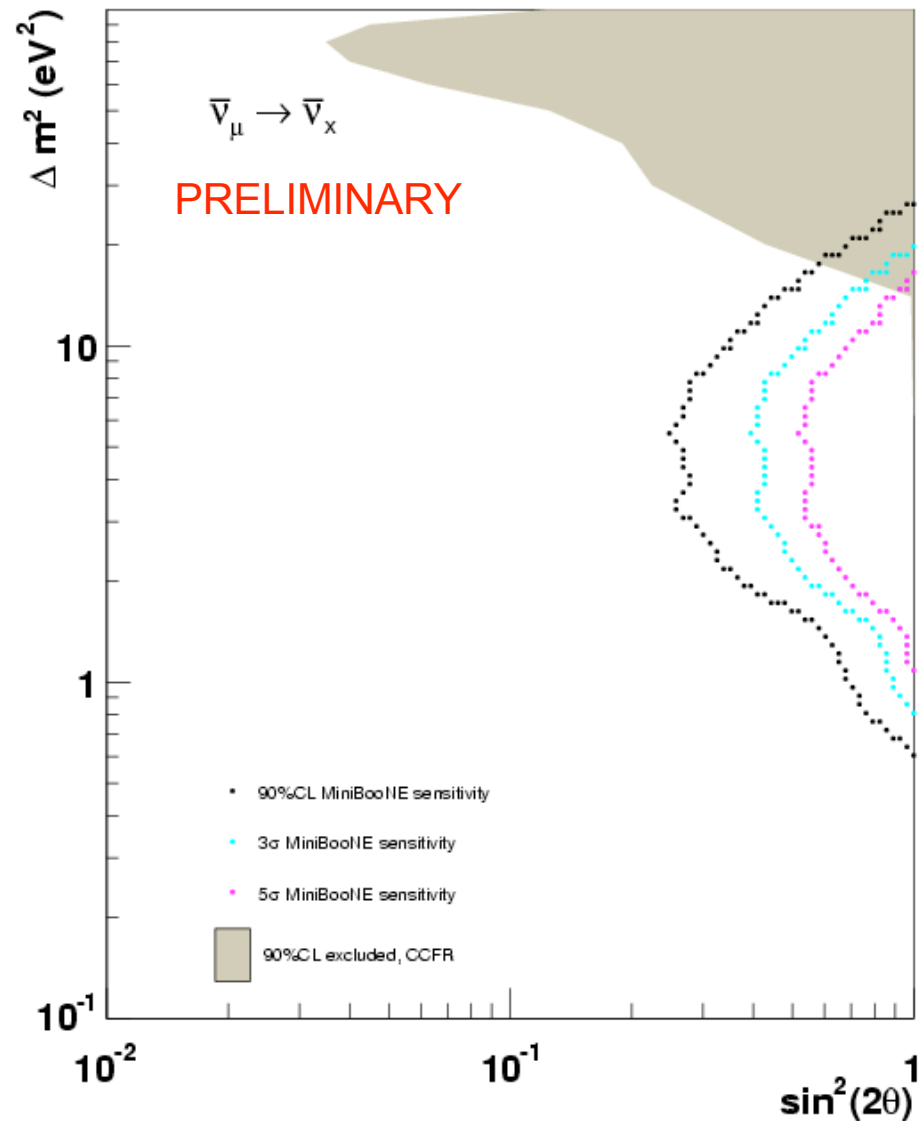
The neutrino and antineutrino spectrums are quite similar

If we change the normalization of the antineutrinos ( $N\pi^-$ ) differently from the neutrinos ( $N\pi^+$ ), the effect on the shape of the antineutrino sample is less than the size of the statistical errors



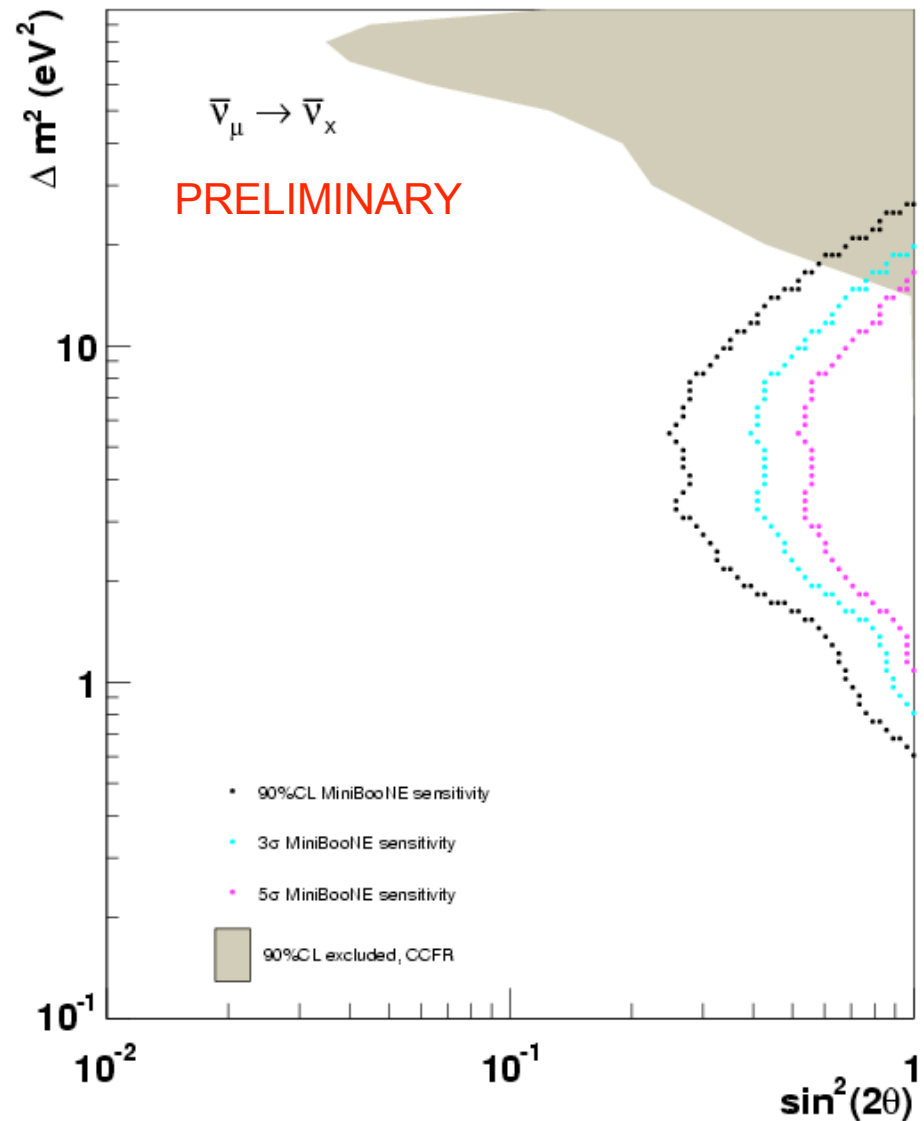
# Antineutrino disappearance sensitivity

- 90% CL antineutrino disappearance sensitivity for  $3.38\text{E}20$  POT
- Plot assumes no  $\nu_\mu$  disappearance based on prior work
- Substantial new parameter space covered

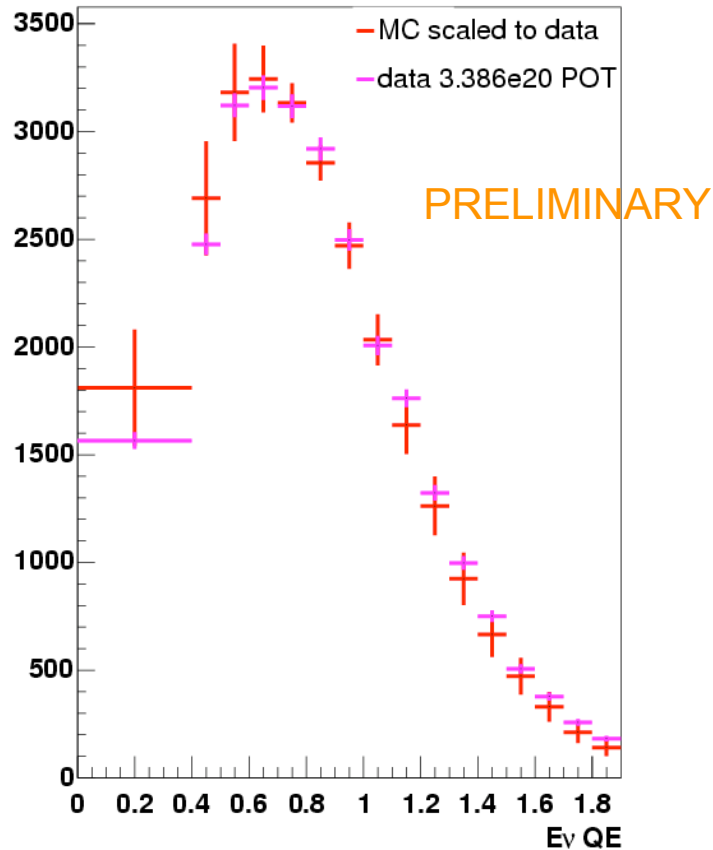


# Antineutrino disappearance sensitivity

- 90% CL antineutrino disappearance sensitivity for  $3.38\text{E}20$  POT
- Plot assumes no  $\nu_\mu$  disappearance based on prior work
- Substantial new parameter space covered
- How about data?



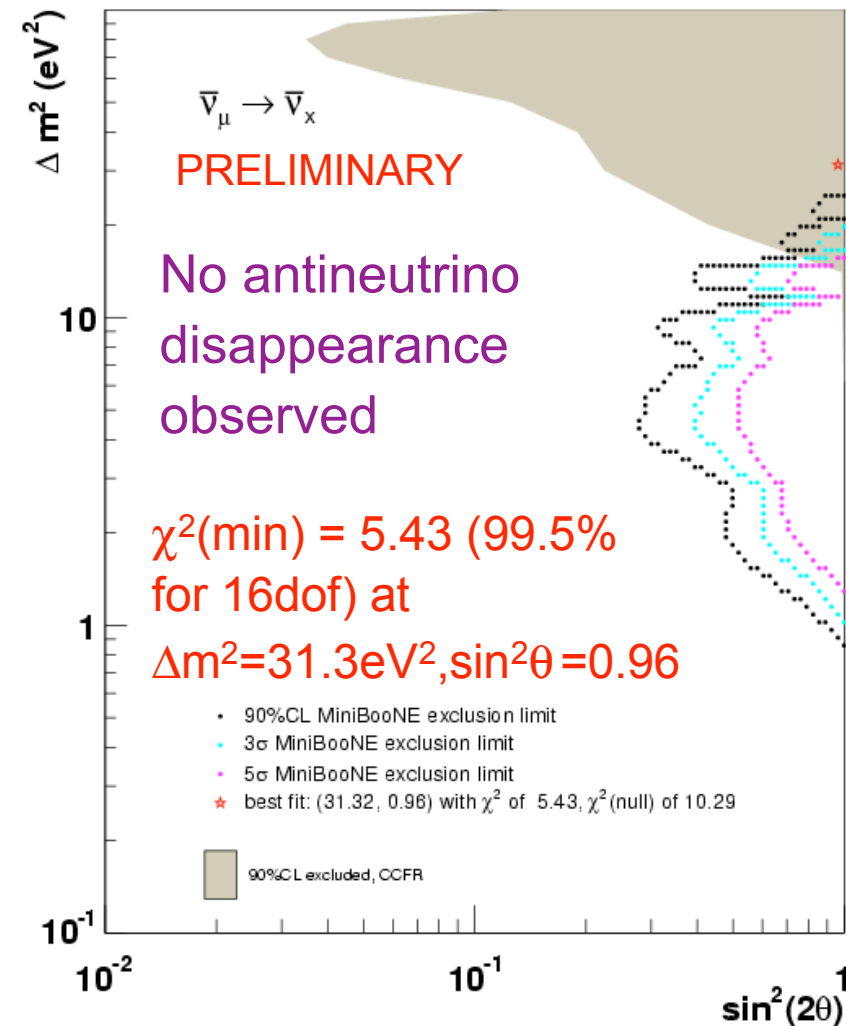
# Antineutrino disappearance results



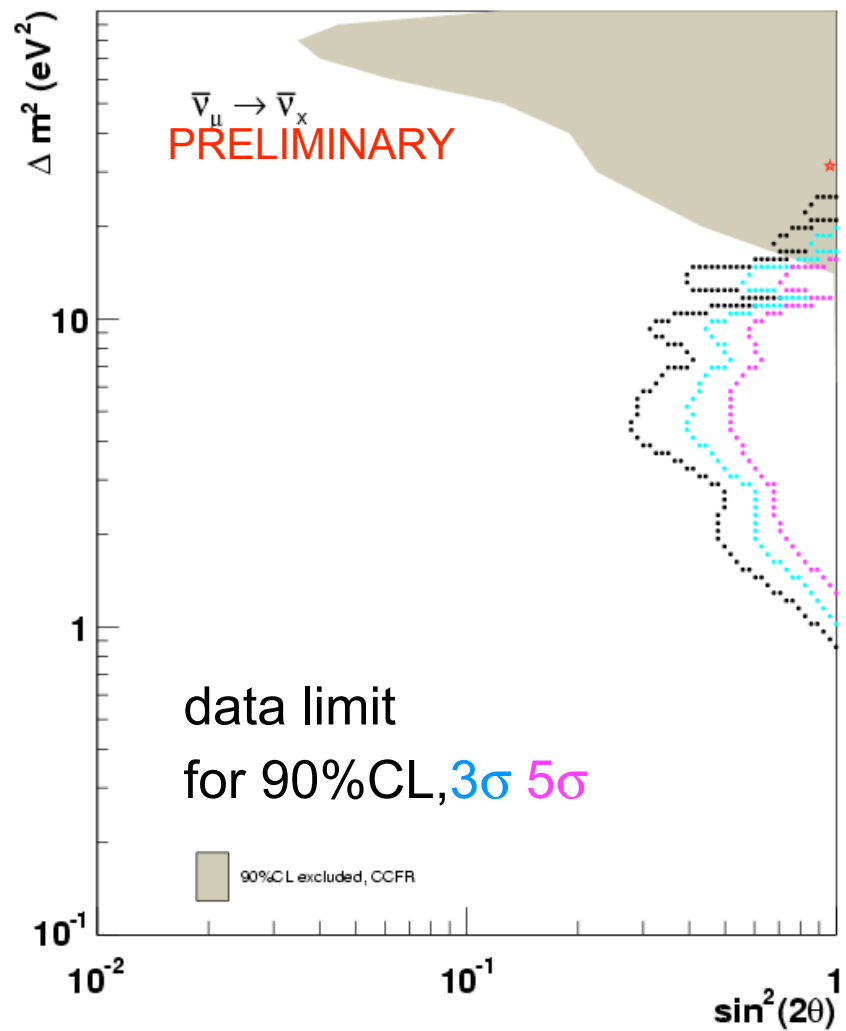
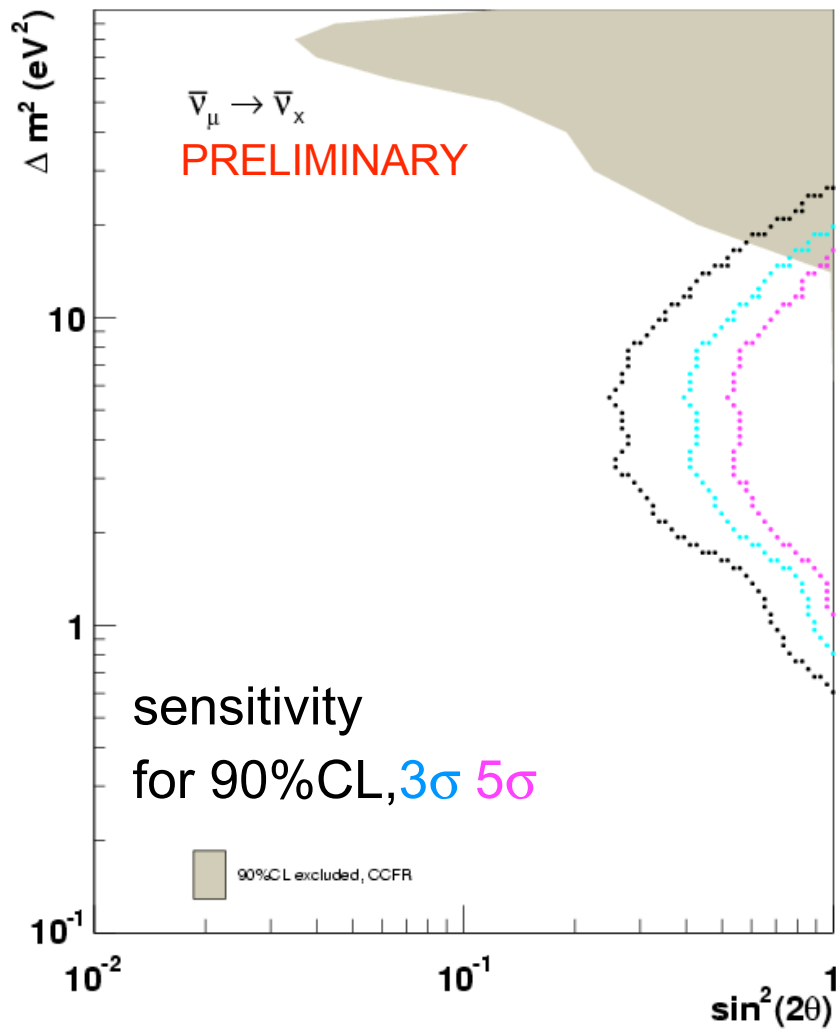
3.38e20 dataset w/ statistical errors  
null oscillation w/ diagonal shape errors

$$\chi^2(\text{null}) = 10.29 \text{ (85\% for 16dof)}$$

$$\chi^2(\text{null, stat only}) = 109 \text{ (16dof)}$$



# Antineutrino disappearance results

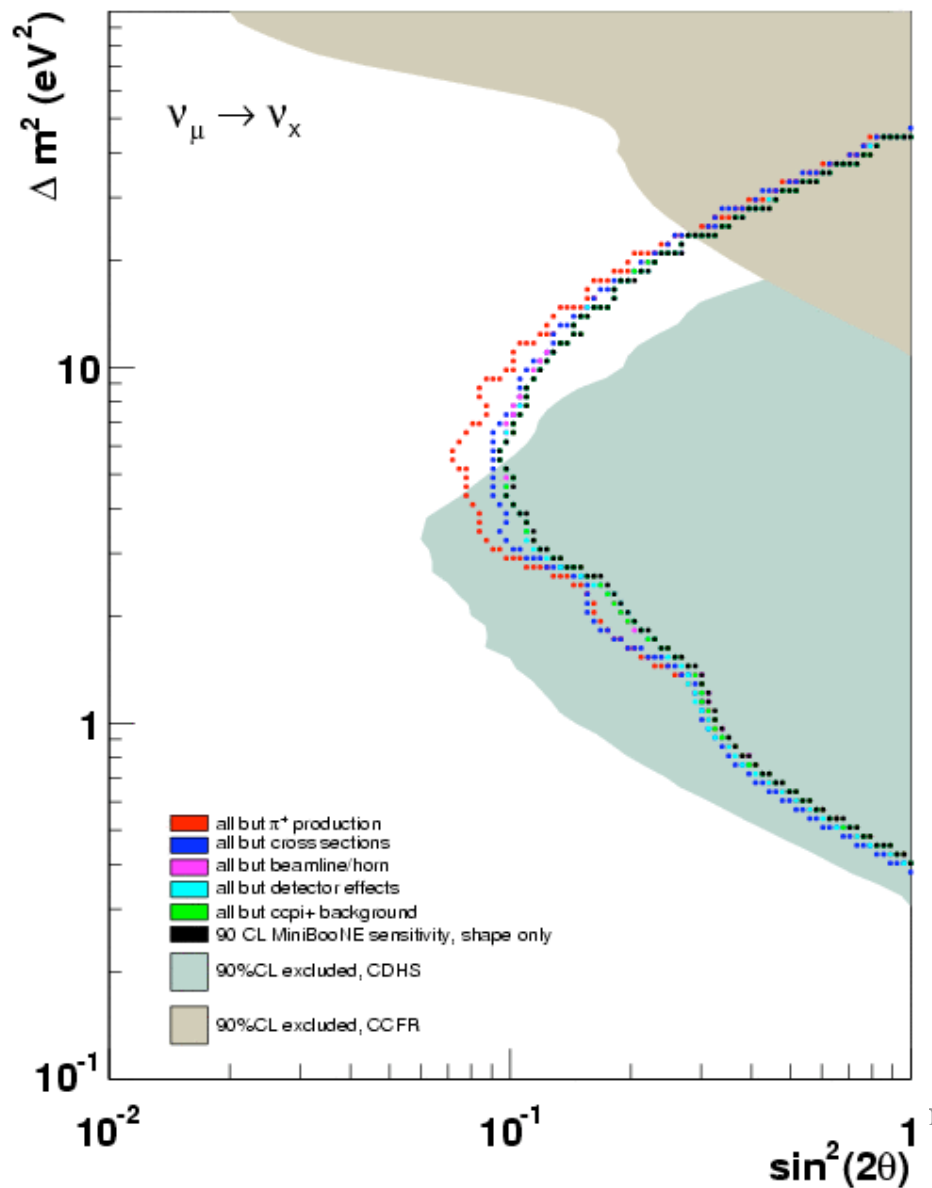


# Overview

- 1) Neutrino oscillation
- 2) MiniBooNE experiment
- 3) MiniBooNE-only neutrino disappearance analysis
- 4) Antineutrino disappearance analysis
- 5) Improvements to disappearance analysis
- 6) Conclusion



# Improvements to $\nu_\mu$ disappearance?



Remove each source of error one at a time, which error affects 90% shape only sensitivity most?

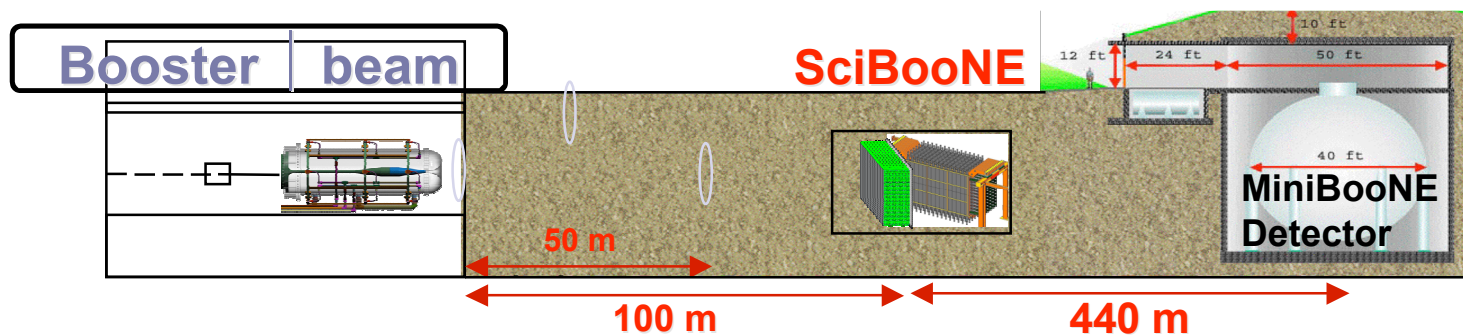
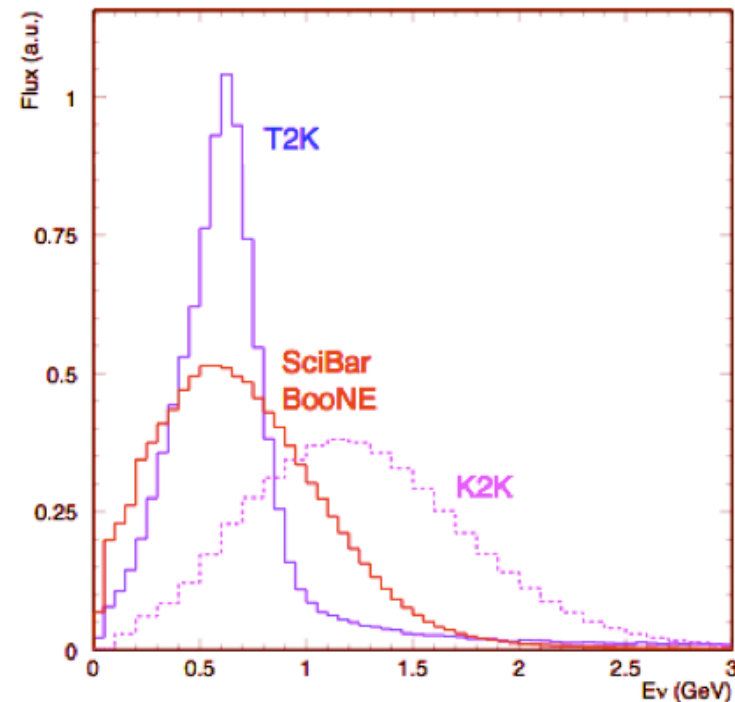
Dominant errors are **flux** and **cross section**

Near detector constrains both

Incorporate SciBooNE data!

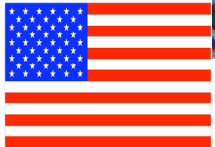
# SciBooNE

- Insert preexisting (free!) fine grained tracking detectors into Booster Neutrino Beamline
- Provide cross section information for future oscillation experiments, such as T2K
  - Similar energy range
- Also provides a near detector for MiniBooNE
  - Nearly identical flux, identical target (carbon)



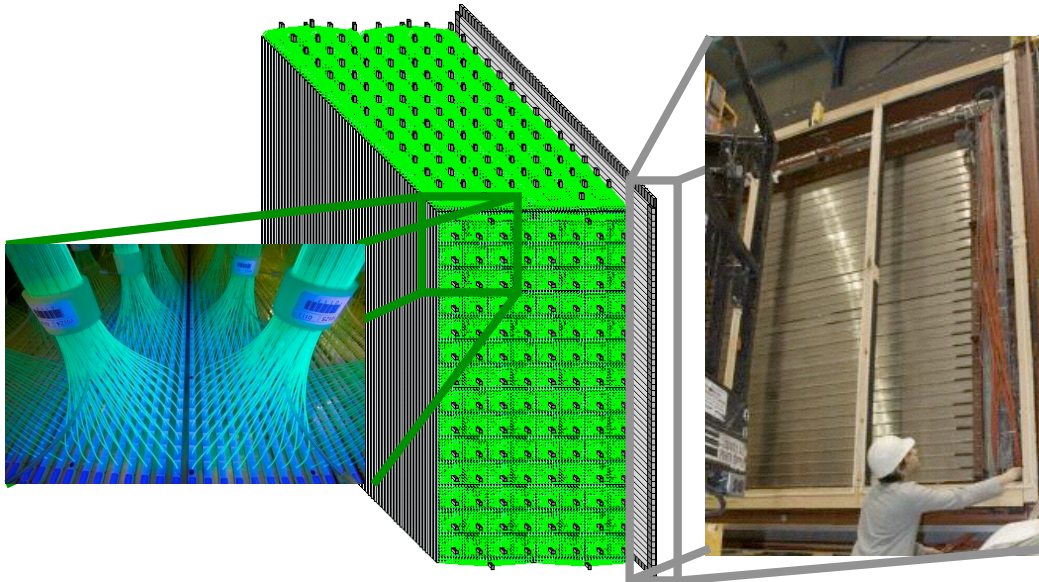
# The SciBooNE collaboration

*A selection of SciBooNE collaborators at the last collaboration meeting. March 2008*



*Universitat Autònoma de Barcelona  
University of Cincinnati  
University of Colorado  
Columbia University  
Fermi National Accelerator Laboratory  
High Energy Accelerator Research  
Organization (KEK)  
Imperial College London  
Indiana University  
Institute for Cosmic Ray Research  
Kyoto University  
Los Alamos National Laboratory  
Louisiana State University  
Purdue University Calumet  
Università degli Studi di Roma  
and INFN-Roma  
Saint Mary's University of Minnesota  
Tokyo Institute of Technology  
Universidad de Valencia*

# SciBooNE detectors



- SciBar vertex detector

- 14,336 channel extruded scintillator with WLS fibers
  - 64 channel Multi-Anode PMT readout

- Used in K2K experiment

- Electron Calorimeter (EC)

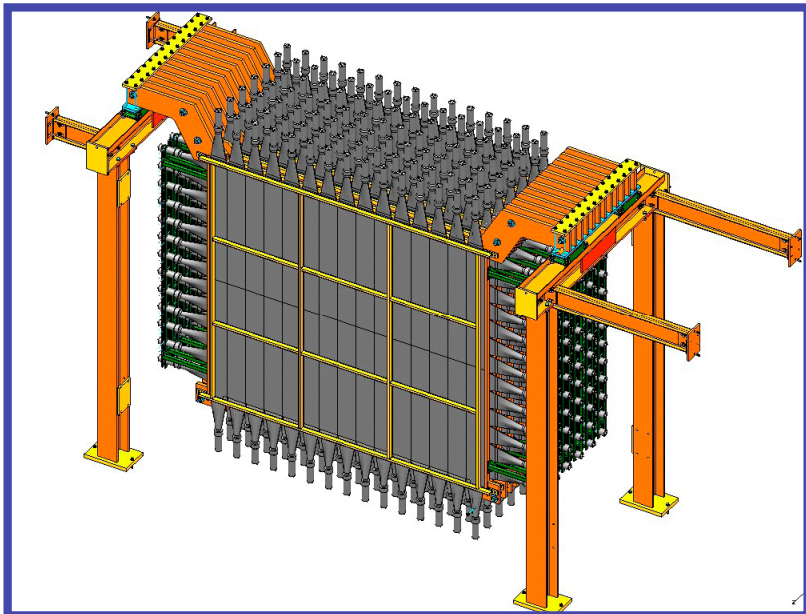
- 256 channel “spaghetti” calorimeter (scintillating fiber & lead foil)

- Used in CHORUS, later K2K

- Muon range detector (MRD)

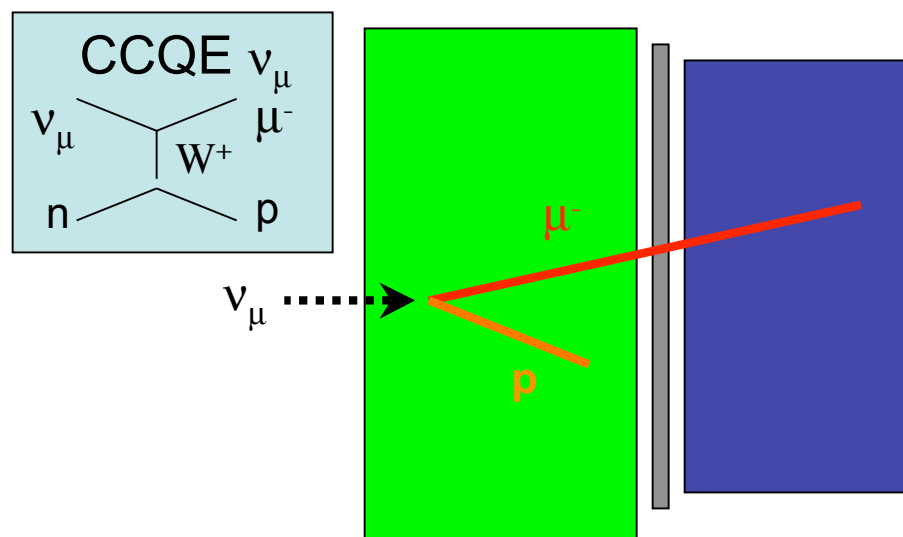
- 362 scintillator counters strapped to 12 iron planes

- Built at FNAL with spare parts



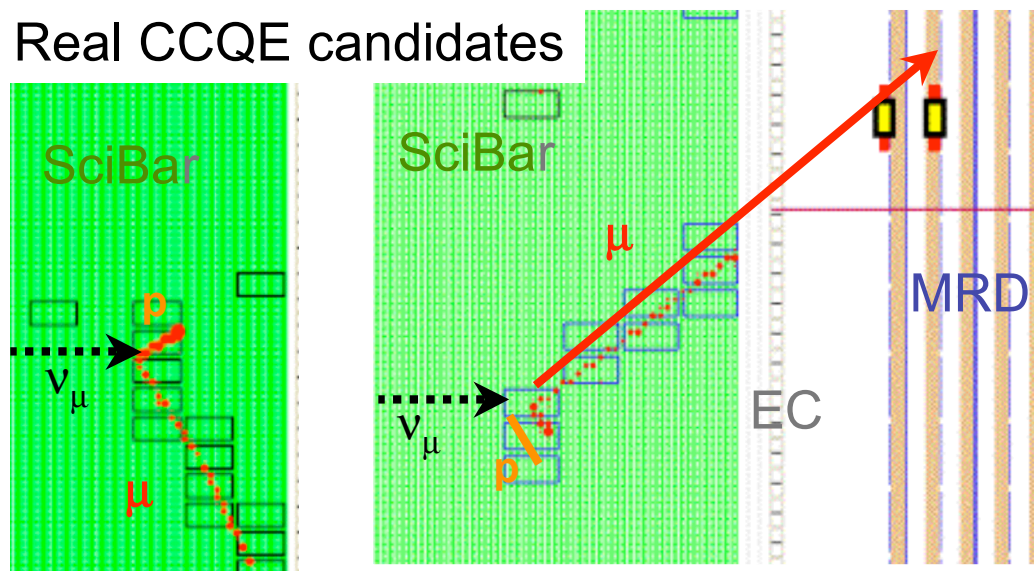


# SciBooNE detectors

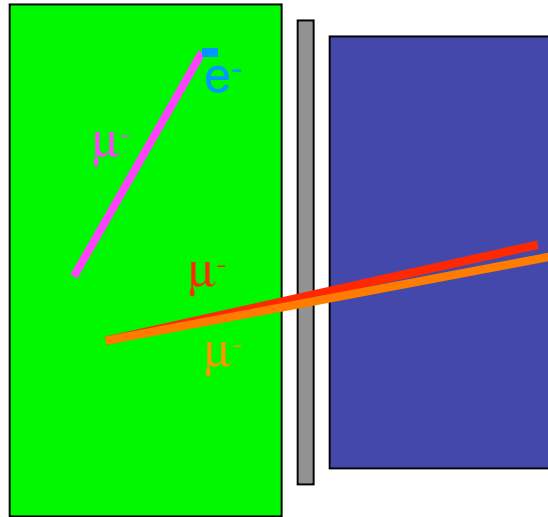


- **SciBar vertex detector**  
tracks  $>8\text{cm}$  are reconstructable  
Can use  $dE/dX$  to distinguish  
protons from pions, muons
- **Electron Calorimeter (EC)**  
electron/muon separation  
 $11X^0$ ,  $14\% \sqrt{E}$
- **Muon range detector (MRD)**  
Measures muons  $< 1.2 \text{ GeV}$   
to  $\sim 10\%$  resolution

Real CCQE candidates

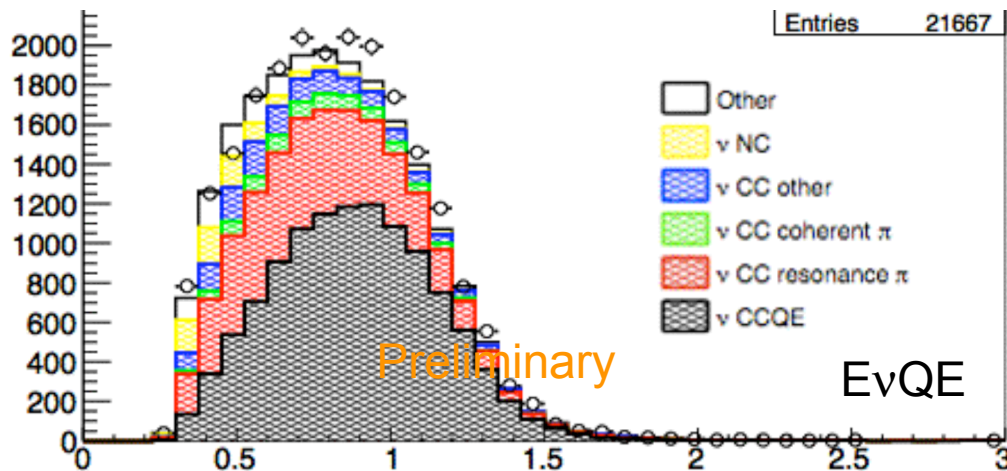


# SciBooNE data samples



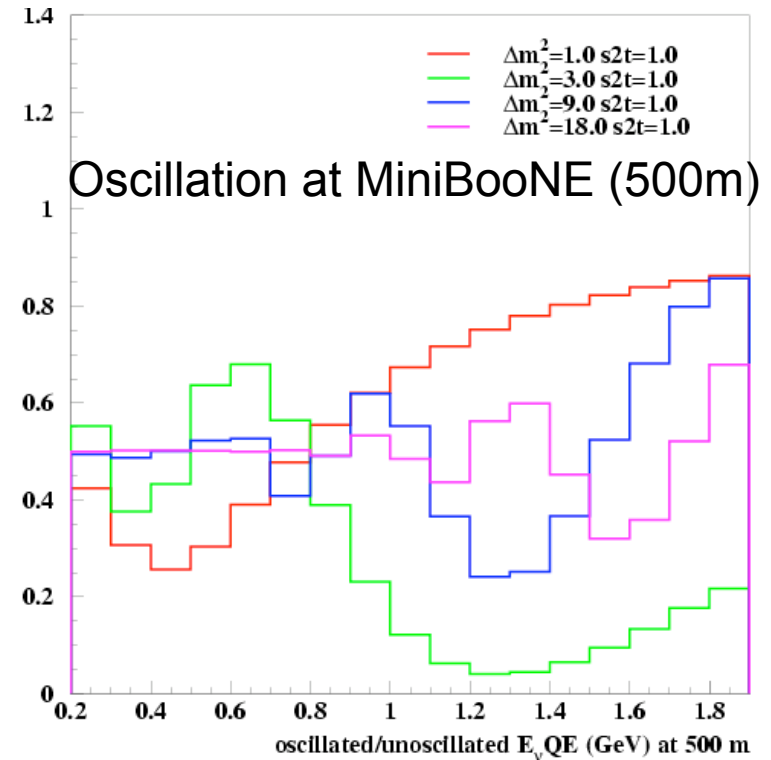
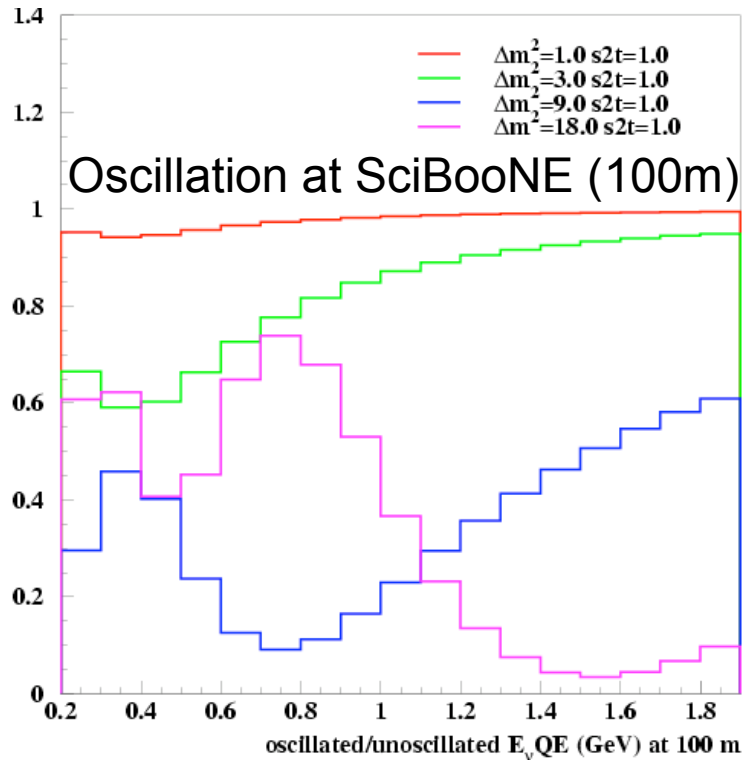
- Tag CCQE events within SciBar using decay electron  
“SciBar contained”
- Tag CC events with muon in MRD  
MRD Matched → “MRD Stopped” or “MRD Penetrated”

“MRD Stopped” sample



- Already developing data sets  
neutrino mode:  $0.99e20$  POT  
~30k MRD Matched events  
antineutrino mode:  $1.53e20$  POT  
~13k MRD Matched events

# Joint MiniBooNE/SciBooNE analysis



Fit will be able to include normalization information from SciBooNE

For some oscillation signals, oscillation can be seen in SciBooNE

The flux and cross section will cancel, but the amount of correlation between the two detectors is reduced by statistics and detector errors

# Conclusion

- MiniBooNE observes no neutrino or antineutrino disappearance

Will add constraints to 3+N models

Limits CPT violating models

- Future work will include SciBooNE as a near detector constraint on the disappearance analysis

- Additional BooNE news:

SciBooNE has finished its first result on  $CC\pi^+$  coherent production

hep-ex/0811.0369 on archive as of this Monday

A host of MiniBooNE cross section analyses are also in the works

- $CC\pi^+/CCQE$  ratio measurement
- NC  $\pi^0$  coherent/resonant fraction for antineutrino events
- Differential cross sections ( $CCQE$ , NC elastic,  $NC\pi^0$ ,  $CC\pi^+$ ,  $CC\pi^0$ )

Direct LSND test, electron antineutrino appearance results in December